

# SIMP Spectroscopy

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**ABSTRACT:** We study the interactions between strongly interacting massive particle dark matter and the Standard Model via a massive vector boson that is kinetically mixed with the hypercharge gauge boson. The relic abundance is set by  $3 \rightarrow 2$  self-interactions of the dark matter, while the interactions with the vector mediator enable kinetic equilibrium between the dark and visible sectors. We show that a wide range of parameters is phenomenologically viable and can be probed in various ways. Astrophysical and cosmological constraints are evaded due to the  $p$ -wave nature of dark matter annihilation into visible particles, while direct detection methods using electron recoils can be sensitive to parts of the parameter space. In addition, we propose performing spectroscopy of the strongly coupled dark sector at  $e^+e^-$  colliders, where the energy of a mono-photon can track the resonance structure of the dark sector. Alternatively, some resonances may decay back into Standard Model leptons or jets, realizing ‘hidden valley’ phenomenology at the LHC and ILC in a concrete fashion.

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## 1 Introduction

The weakly interacting massive particle (WIMP) has long been the dominant paradigm for explaining the presence of dark matter. This is because of the so-called ‘WIMP miracle’—that weak-scale dark matter with weak interactions gives the correct relic abundance if the dark matter annihilates to Standard Model (SM) particles—and that we expect new weak-scale particles to exist and address the hierarchy problem. There is, however, no compelling evidence of the WIMP’s existence, and new particles at the weak-scale have yet to be discovered. As a result, considerable theoretical and experimental effort has been put into exploring dark-matter scenarios beyond the WIMP.

One recent proposal to explain the particle nature of dark matter is the strongly interacting massive particle (SIMP) mechanism [1]. SIMPs have several distinguishing characteristics compared to WIMPs. First, the process responsible for the dark matter relic abundance is a  $3 \rightarrow 2$  self-annihilation process consisting entirely of dark matter particles. This mechanism of thermal freezeout predicts dark matter that is lighter than the WIMP, with masses typically between an MeV and a GeV. Indeed, QCD-scale dark matter with strong self-couplings gives the correct relic abundance. The strong self-interactions of the SIMP also lead to strong self-scatterings, which can address longstanding puzzles in small-scale structure. The  $3 \rightarrow 2$  process pumps heat into the system [2], and so the dark matter must dump its entropy, in order to stay cool during freezeout. This can be done by interacting either with new light degrees of freedom or with abundant light SM particles. In the latter case, couplings to the SM of a particular size are necessary for maintaining thermal equilibrium with the thermal bath, leading to distinct experimental dark-matter signatures.

In Ref. [3] it was proposed that SIMPs could be the pions (pseudo-Nambu-Goldstone bosons) of a strongly coupled QCD-like confining hidden sector, where the Wess-Zumino-Witten term [4–6] generated the  $3 \rightarrow 2$  number-changing process. Requiring the validity of chiral perturbation theory, together with dark matter self-interaction constraints, suggests that the mass of the pion and its decay constant should be of order a few  $\times 100$  MeV, resembling the pions and kaons of the SM.

Various aspects of SIMP dark matter are being studied in the literature (see, *e.g.*, Refs. [7–13]). In Ref. [8], a particular mediator between the SM and SIMPs was studied, towards maintaining thermal contact between the two sectors. The setup was of a  $U(1)_D$  symmetry under which the pions are charged, and whose gauge boson kinetically mixes with the hypercharge gauge boson of the SM.

In this work, we explore in further detail the prospects for kinetically mixed vector mediation between the SIMP and SM sectors. We include the possibility that the dark matter belongs to various different confining gauge groups, finding appropriate subgroups of the residual global symmetries that can be gauged. We find the experimental constraints on the kinetically mixed parameter space, and give future experimental probes of these interactions.

Importantly, we propose performing *spectroscopy* of the strongly coupled sector at  $e^+e^-$  colliders, where the energy of a mono-photon, in  $e^+e^- \rightarrow \gamma + \text{inv}$ , can be used as a

tracer for the resonance structure of the SIMP sector. More generally, dark spectroscopy can be done for any strongly coupled dark sector that couples (in a conceptually similar way) to the SM.

The outline of this paper is as follows. Section 2 reviews the pion-setup of the SIMP mechanism, and Section 3 sets up the kinetically mixed hidden photon. In Section 4 we describe the dark photon as a mediator between SIMPs and the SM: this includes the embedding of the dark photon into the SIMP setup for different confining gauge groups, as well as thermalization and annihilation. Section 5 details the decays of the dark sector, visible and invisible. We describe our proposal of dark spectroscopy in Section 6, discuss future prospects in Section 7, and conclude in Section 8.

## 2 SIMP review

We begin by reviewing the SIMP setup presented in Ref. [3], where the dark matter is a pion (pseudo-Nambu-Goldstone boson) of a confining hidden sector, with the Wess-Zumino-Witten (WZW) action generating the  $3 \rightarrow 2$  process.

We consider an  $SU(N_c)$  gauge theory with  $N_f$  Dirac fermions, an  $SO(N_c)$  gauge theory with  $N_f$  Weyl fermions, or an  $Sp(N_c)$  gauge theory with  $2N_f$  Weyl-fermions,  $q_i$ , in the  $N_c$ -dimensional representation of the gauge group, with  $N_c$  the number of colors ( $N_c$  is even for  $Sp(N_c)$ ). Neglecting quark masses, the Lagrangian takes the simple form

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}^a F^{\mu\nu a} + q_i^\dagger i \bar{\sigma}^\mu D_\mu q_i. \quad (2.1)$$

The Lagrangian contains an exact global-symmetry,  $G$ , which with a sufficiently small  $N_f$  for a given  $N_c$ , leads to chiral symmetry breaking of the global symmetry down to a subgroup  $H$ . The expected symmetry-breaking pattern  $G/H$  is  $SU(N_f) \times SU(N_f)/SU(N_f)$ ,  $SU(N_f)/SO(N_f)$ , and  $SU(2N_f)/Sp(2N_f)$ , for an  $SU(N_c)$ ,  $SO(N_c)$ , and  $Sp(N_c)$  gauge theory, respectively.

The low-energy effective chiral action for the above theories is

$$S_{\text{eff}} = \int d^4x \frac{f_\pi^2}{16} \text{Tr} \partial_\mu \Sigma \partial^\mu \Sigma^\dagger + \Gamma_{\text{WZW}}, \quad (2.2)$$

where  $\Sigma$  parameterizes the coset space  $G/H$ ,

$$\Sigma = \begin{cases} \exp(2i\pi/f_\pi)I & \text{for } SU(N_c) \text{ or } SO(N_c), \\ \exp(2i\pi/f_\pi)J & \text{for } Sp(N_c). \end{cases} \quad (2.3)$$

Here  $\pi = \pi_a T_a$ ,  $T_a$  are the broken generators of  $G/H$  and  $J = i\sigma_2 \otimes I_{N_f}$  is the  $Sp(2N_f)$  invariant. We use the normalization  $\text{Tr} T^a T^b = 2\delta^{ab}$ .

The ungauged Wess-Zumino-Witten action,  $\Gamma_{\text{WZW}}$ , takes the form:

$$\Gamma_{\text{WZW}} = \frac{-iN_c}{240\pi^2} \int_{M^5} \text{Tr}(d\Sigma \Sigma^{-1})^5, \quad (2.4)$$

where the integral is over a five-dimensional manifold ( $M^5$ ) whose boundary is identified with the ordinary four-dimensional Minkowski space. The WZW-action is non-vanishing

provided that the fifth homotopy group of the coset space,  $\pi_5(G/H)$ , is non-trivial. This is satisfied when  $N_f \geq 3$  for  $SU(N_c)$  and  $SO(N_c)$  gauge theories and  $N_f \geq 2$  for  $Sp(N_c)$  gauge theories.

Expanding the effective-action to leading order in the pion fields, the interaction Lagrangian is

$$\mathcal{L}_{\text{int}} = -\frac{1}{6f_\pi^2} \text{Tr} (\pi^2 \partial^\mu \pi \partial_\mu \pi - \pi \partial^\mu \pi \pi \partial_\mu \pi) + \frac{2N_c}{15\pi^2 f_\pi^5} \epsilon^{\mu\nu\rho\sigma} \text{Tr} (\pi \partial_\mu \pi \partial_\nu \pi \partial_\rho \pi \partial_\sigma \pi) + \mathcal{O}(\pi^6). \quad (2.5)$$

The 5-point interactions coming from the WZW-action, which involve five different pion fields, enable the  $3 \rightarrow 2$  process of the SIMP mechanism.

The SIMPlest model is that of the smallest particle content, which is the  $Sp(2) \simeq SU(2)$  gauge group with  $N_f = 2$ , where the coset space is  $SU(4)/Sp(4) = SO(6)/SO(5) = S^5$ , with  $\pi_5(S^5) = \mathbb{Z}$ . The five pions are the minimum number of pions needed for the  $3 \rightarrow 2$  process via the WZW term. In what follows, this model will often be used as an example to demonstrate the phenomenology of the proposal here.

Without an explicit mass term for the quarks, the pions are exact Goldstone-bosons, and therefore massless. For the gauge theories considered here, an  $H$ -invariant quark mass can be added to the Lagrangian,

$$\mathcal{L}_{\text{mass}} = \begin{cases} m_q I^{ij} \bar{q}_i q_j + h.c., & \text{for } SU(N_c) \\ \frac{1}{2} m_q I^{ij} q_i q_j + h.c., & \text{for } SO(N_c) \\ \frac{1}{2} m_q J^{ij} q_i q_j + h.c., & \text{for } Sp(N_c) \end{cases} \quad (2.6)$$

which induces a mass for all the pions,

$$\begin{aligned} \mathcal{L}_{\text{eff-mass}} &= \begin{cases} -4f_\pi^2 m_\pi^2 \text{Tr} \Sigma + h.c. & \text{for } SU(N_c), SO(N_c) \\ -4f_\pi^2 m_\pi^2 \text{Tr} J \Sigma + h.c. & \text{for } Sp(N_c) \end{cases} \\ &= -\frac{m_\pi^2}{4} \text{Tr} \pi^2 + \frac{m_\pi^2}{12f_\pi^2} \text{Tr} \pi^4 + \mathcal{O}(\pi^6). \end{aligned} \quad (2.7)$$

The  $H$ -invariant mass-term ensures that the chiral Lagrangian respects the unbroken symmetry. The pions transform as a non-trivial representation under  $H$  and thus are stable, as they are the lightest fields with non-trivial quantum numbers. In what follows, we will always assume a mass-term for the quarks that preserves some subgroup of the flavor symmetry  $H$ , which acts to stabilize the pions.

The thermally averaged  $3 \rightarrow 2$  annihilation cross-section can be calculated from the effective 5-point interaction in Eq. (2.5),

$$\langle \sigma v^2 \rangle_{3 \rightarrow 2} = \frac{5\sqrt{5}}{2\pi^5 x_f^2} \frac{N_c^2 m_\pi^5}{f_\pi^{10}} \frac{t^2}{N_\pi^3}, \quad (2.8)$$

where  $x_f = m_\pi/T_f \simeq 20$  and  $T_f$  is the bath temperature at freeze-out. The factor of  $t^2/N_\pi^3$  is combinatorial factor that depends on  $N_f$  and the choice of the confining gauge-theory; further details are found in the Appendix of Ref. [3]. For a given  $N_f$  and  $N_c$ , obtaining

the correct relic abundance via  $3 \rightarrow 2$  annihilation gives an explicit relationship between  $m_\pi$  and  $f_\pi$ .

There are strong constraints on large dark-matter self scatterings from bullet-cluster observations [14–16] and halo shape simulations [17, 18], which give roughly  $\sigma_{\text{scatter}}/m_{\text{DM}} \lesssim 1 \text{ cm}^2/\text{g}$ . In terms of the pions, the self-scattering cross section can be calculated to leading order using the 4-point interactions of the chiral Lagrangian [Eqs. (2.5) and (2.7)],

$$\sigma_{\text{scatter}} = \frac{m_\pi^2}{32\pi f_\pi^4} \frac{a^2}{N_\pi^2}, \quad (2.9)$$

where the factor of  $a^2/N_\pi^2$  is a combinatorial factor that depends on  $N_f$  and the choice of the confining gauge-theory; further details can be found in the Appendix of Ref. [3].

Applying the constraints on self-interactions, that the pions produce the observed dark-matter relic abundance, and assuming validity of chiral perturbation theory  $m_\pi/f_\pi \lesssim 2\pi$ , points to a preferred region of parameter space which is very similar to that of QCD:

$$m_\pi \sim 300 \text{ MeV}, \quad f_\pi \sim \text{few} \times m_\pi. \quad (2.10)$$

This corresponds to the strongly interacting regime of the theory, where the strong dynamics can induce  $\mathcal{O}(1)$  changes to the above; Eq. (2.10) should be thought of as a mere proxy to the scales involved.

We note that the self-scattering cross sections predicted by Eq. (2.9) are of the right size to reconcile discrepancies in N-body simulations with the observed small-scale structure, such as the ‘core vs. cusp’ and ‘two big to fail’ puzzles [17–24]. Additionally, a recent study of the Abell 3827 galaxy cluster showed evidence for self-interactions of dark matter [25, 26]. Ref. [25] claimed the self-scattering cross section should be  $\sigma_{\text{scatter}}/m = (1.7 \pm 0.7) \times 10^{-4} \text{ cm}^2\text{g}^{-1} \times (t_{\text{infall}}/10^9 \text{ yr})^{-2}$ , but a reanalysis in Ref. [26] finds  $\sigma_{\text{scatter}}/m = 1.5 \text{ cm}^2\text{g}^{-1}$  as the preferred value, similar to the rate needed to address the small-scale structure puzzles and the rate expected from the SIMP mechanism.

### 3 Dark photon review

We consider a massive dark photon, which is the  $U(1)_D$  gauge boson, that is kinetically mixed with the hypercharge gauge boson. Extensively studied in the literature, here we summarize the relevant parts of the dark photon Lagrangian as well as relevant experimental limits, including several new limits.

#### 3.1 Kinetic mixing

We take a  $U(1)_D$  gauge theory with a gauge field  $\mathcal{A}_\mu$  which has kinetic mixing with the  $U(1)_Y$  gauge field  $B_\mu$ ,

$$\mathcal{L}_\mathcal{A} = -\frac{1}{4}\mathcal{A}_{\mu\nu}\mathcal{A}^{\mu\nu} - \frac{\sin\chi}{2}B_{\mu\nu}\mathcal{A}^{\mu\nu} + \frac{1}{2}m_V^2\mathcal{A}_\mu\mathcal{A}^\mu. \quad (3.1)$$

The gauge boson’s mass may arise either from a hidden-photon Higgs mechanism or the Stückelberg trick.

After electroweak symmetry breaking occurs,  $\mathcal{A}_\mu$  becomes a mixture of the  $Z$ -boson and the dark photon  $V$  [8, 27],

$$V^\mu = -\frac{s_\zeta}{c_\chi} Z^\mu + \frac{c_\zeta}{c_\chi} \mathcal{A}^\mu, \quad (3.2)$$

where

$$\tan 2\zeta = \frac{m_Z^2 s_W \sin 2\chi}{m_V^2 - m_Z^2 (c_\chi^2 - s_W^2 s_\chi^2)}. \quad (3.3)$$

The mass eigenvalues are then

$$m_\pm^2 = \frac{1}{2} \left[ m_Z^2 (1 + s_W^2 t_\chi^2) + \frac{m_V^2}{c_\chi^2} \pm \sqrt{\left( m_Z^2 (1 + s_W^2 t_\chi^2) + \frac{m_V^2}{c_\chi^2} \right)^2 - \frac{4}{c_\chi^2} m_Z^2 m_V^2} \right]. \quad (3.4)$$

The vector bosons couple to the current operators as

$$\begin{aligned} \mathcal{L}_D = & A_\mu J_{EM}^\mu + Z_\mu \left[ c_W s_\zeta t_\chi J_{EM}^\mu + (c_\zeta - s_W t_\chi s_\zeta) J_Z^\mu - \frac{s_\zeta}{c_\chi} J_D^\mu \right] \\ & + V_\mu \left[ J_{EM}^\mu (-c_W c_\zeta t_\chi) + J_Z^\mu (s_\zeta + s_W t_\chi c_\zeta) + \frac{c_\zeta}{c_\chi} J_D^\mu \right], \end{aligned} \quad (3.5)$$

For later convenience, we define the coupling strength of the dark photon to the electromagnetic and  $Z$  currents,

$$\epsilon_\gamma \equiv -c_W c_\zeta t_\chi \quad \text{and} \quad \epsilon_Z \equiv s_\zeta + s_W t_\chi c_\zeta. \quad (3.6)$$

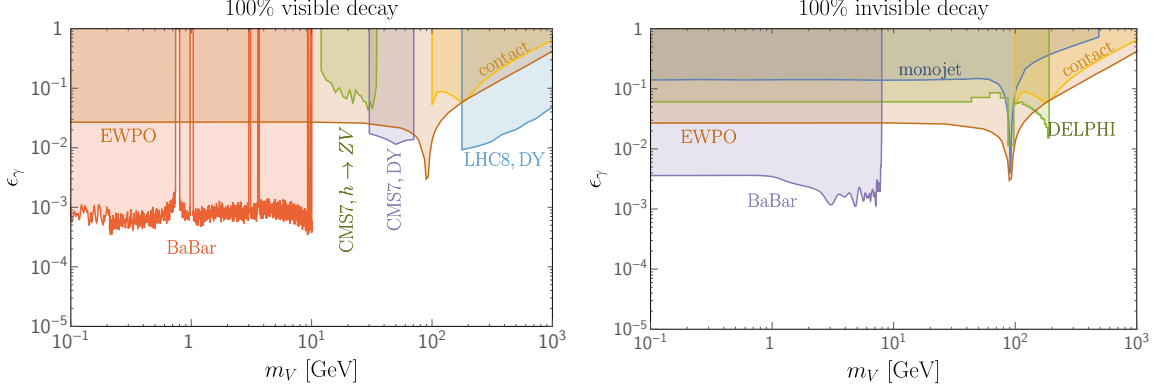
### 3.2 Experimental limits on dark photons

In order to understand the range of parameters viable for the SIMP mechanism, we now assemble the current experimental limits on dark photons. There are two extreme cases of the dark photon decays to be considered: (I) 100% branching fraction into Standard Model particles (left panel of Fig. 1), and (II) 100% branching fraction into invisible particles (right panel of Fig. 1).

Most of the limits presented in Fig. 1 are reproduced from Refs. [28–30]. However, a few comments are in order. First, we do not consider the case of a dark photon lighter than twice the dark pion. The reason is that if the coupling of this process is strong enough to achieve kinetic equilibrium between the two sectors at the time of the freeze-out, it would cause the annihilation rate of  $\pi\pi \rightarrow \pi V$  to dominate over the  $\pi\pi\pi \rightarrow \pi\pi$  process required for the SIMP mechanism [8]. The constraint from electroweak precision observables (labeled **EWPO**, brown) [29] is independent of how the dark photon decays and is shown in both panels of Fig. 1. Another constraint relevant to both visible and invisible decays comes from the search for a contact four-fermion operator at LEP II [31] (labeled **contact**, orange), which was not discussed in Refs. [28–30]. This process is nominally weaker than the limit from EWPO. We include it, however, since it is expected to improve at the ILC by roughly an order of magnitude [32] (see Fig. 7).

Bounds from visible decays<sup>1</sup> of a dark photon are shown in the left panel of Fig. 1. Starting from the low  $m_V$  region, the first constraint (labeled **BaBar**, red) is from a

<sup>1</sup>We define ‘visible decays’ to include decays into neutrinos in this paper.



**Figure 1.** Constraints on  $\epsilon_\gamma$  as a function of  $m_V$  from dark photon decays; see text for details. **Left:** 100% visible decays (including decays to  $\nu\bar{\nu}$ ); **Right:** 100 % invisible decays.

dedicated search for  $e^+e^- \rightarrow \gamma V$  followed by  $V \rightarrow e^+e^-, \mu^+\mu^-$  at BaBar [33]. The next constraint (labeled  $\text{CMS7}, h \rightarrow ZV$ , green) is the reinterpretation of the study of  $h \rightarrow Z^{(*)}Z^* \rightarrow 4\ell$  by CMS [34] as a search for  $h \rightarrow Z^{(*)}V \rightarrow 4\ell$  [30]. The third bound (labeled  $\text{CMS7}, \text{DY}$ , purple) is from the search for Drell–Yan production of lepton pairs performed by CMS in Ref. [35]. The right most bound (labeled  $\text{LHC8}, \text{DY}$ , blue) also comes from the search for Drell–Yan production of lepton pairs in Ref. [36], using the ATLAS data [37].

Similarly, there are bounds specific for the invisible decays of the dark photon; these are shown in the right panel of Fig. 1. Starting from the low  $m_V$  region, the first one (labeled  $\text{BaBar}$ , purple) is from a search at BaBar for  $\Upsilon(3S) \rightarrow \gamma A^0$ , where  $A^0$  is a pseudo-scalar boson that decays invisibly [38]. Note that  $\Upsilon(3S) \rightarrow \gamma V$  is forbidden by charge conjugation invariance, but the limit can be re-interpreted as that on the continuum production  $e^+e^- \rightarrow \gamma V$  followed by  $V \rightarrow \text{invisible}$  [28, 29]. The next bound (labeled  $\text{monojet}$ , blue) is for  $pp \rightarrow \text{jet} + V$ ,  $V \rightarrow \text{invisible}$ . Here the limit is scaled from the coupling to the up-quark in Ref. [39] that used data from CDF and ATLAS (7 TeV). This bound is currently subdominant to other constraints, but is expected to be improved with higher luminosity and energy at the LHC. Finally, the limit from the mono-photon search at DELPHI [40] (labeled  $\text{DELPHI}$ , green) is reinterpreted, allowing for  $2\sigma$  fluctuations in each energy bin.

All limits are expected to be improved in the future by the LHC Run-2 and beyond, SuperKEKB, ILC and/or a 100 TeV  $pp$  collider. The limits on a given model depend on the branching fractions of the dark photon into visible particles vs. invisible modes. These branching ratios will be studied in Section 5. Future prospects will be discussed in Section 7.

## 4 The dark photon in SIMP Models

We now study how the dark photon couples to the SIMP sector. We start by identifying the  $U(1)_D$  subgroup of the unbroken global flavor symmetry  $H$ , in the strongly coupled SIMP theory in Section 4.1. We then study two basic constraints for the SIMP setup in



this realization: (I) kinetic equilibrium (thermalization) between the Standard Model and SIMPs during freeze-out (Section 4.2); and (II) dominance of the  $3 \rightarrow 2$  annihilation of SIMPs over the  $2 \rightarrow 2$  annihilation into SM particles,  $\pi\pi \rightarrow V^* \rightarrow \ell^+\ell^-$  (Section 4.3).

#### 4.1 Gauging $U(1)_D$

Here we explore the embedding of the dark photon in the symmetry structure of the SIMP strongly coupled theories. The simplest possibility is to gauge a  $U(1)_D$  part of the unbroken flavor symmetry  $H$  of the coset space  $G/H$ . In principle one can also consider gauging a broken symmetry, in which case one of the pions is eaten by the gauge boson. We consider the case of unbroken generators here only for simplicity. Ref. [8] embedded the  $U(1)_D$  in an  $SU(N_c)$  gauge theory with  $N_f = 3$ ; here we demonstrate the possibility of embedding the  $U(1)_D$  symmetry in general  $Sp(N_c)$ ,  $SU(N_c)$  and  $SO(N_c)$  gauge theories with  $N_f$  flavors.

Which subgroup of  $H$  should be gauged? There are two requirements behind the choice. The first one is that the pions do not decay into dark photons. This can be guaranteed if all the pions transform non-trivially under part of the unbroken flavor symmetry. The other condition is to maintain the near degeneracy of quarks, so that at least five different pions participate in the  $3 \rightarrow 2$  process via the WZW term. The  $U(1)_D$  coupling renormalizes the quark masses,

$$\Delta m_q = \frac{3\alpha_D}{2\pi} Q_q^2 m_q \log \frac{M}{\Lambda}, \quad (4.1)$$

where  $\alpha_D \equiv e_D^2/(4\pi)$  is the dark coupling strength,  $M$  is a UV cutoff such as the Planck scale,  $\Lambda$  is the confining scale, and  $Q_q$  is the  $U(1)_D$ -charge of quark  $q$ . If  $Q_q^2$  is not common to all quarks, quark masses are split at the  $\sim 15\%$  level for  $\alpha_D \approx \alpha$  (with  $\alpha$  the electroweak coupling strength), which may or may not be acceptable. For this reason, and for simplicity, we choose  $U(1)_D$  so that all quarks have the same charge, up to a sign.

We first consider  $Sp(N_c)$  gauge theories, and the resulting flavor coset  $SU(2N_f)/Sp(2N_f)$ . The pions transform in the anti-symmetric tensor of  $Sp(2N_f)$ . This agrees with the intuition that the pions are the  $S$ -wave spin-zero bound state of massive Majorana quarks, with gauge indices contracted by the anti-symmetric symplectic tensor, requiring the anti-symmetric combination of flavor indices. If the quarks are degenerate, the quark mass term also respects  $Sp(2N_f)$ , and hence the degeneracy is preserved by the flavor symmetry. The  $U(1)_D$  is embedded in the global symmetry as

$$SU(N_f) \times U(1)_D \subset Sp(2N_f) \quad (4.2)$$

The quarks transform as  $(\square, +1)$  and  $(\bar{\square}, -1)$  multiplets under  $SU(N_f) \times U(1)_D$ . The pions decompose into representations  $A(\square, +2)$ ,  $A^*(\bar{\square}, -2)$ , and  $\Omega(\text{adj}, 0)$ ,

$$\pi = \begin{pmatrix} \Omega & A \\ -A^* & \Omega^T \end{pmatrix}, \quad (4.3)$$

and since the  $SU(N_f)$  symmetry is exact, neither the anti-symmetric tensor  $A$  nor the adjoint  $\Omega$  can decay, and all pions are stable.

	$U(1)_D \subset H$ embedding	$\pi$ representations
$Sp(N_c)$	$SU(N_f) \times U(1)_D$ $\subset Sp(2N_f)$	$(\square, +2) \oplus (\bar{\square}, -2) \oplus (\text{adj}, 0)$
$SO(N_c)$	$SU(N_f/2) \times U(1)_D$ $\subset SO(N_f)$	$(\square\square, +2) \oplus (\bar{\square}\bar{\square}, -2) \oplus (\text{adj}, 0)$
$SU(N_c)$	$SU(N_1) \times SU(N_2) \times U(1)_D$ $\subset SU(N_f)$	$(\square, \bar{\square}, 2) \oplus (\bar{\square}, \square, -2) \oplus$ $(\text{adj}, 1, 0) \oplus (1, \text{adj}, 0) \oplus (1, 1, 0)$

**Table 1.** Gauging the  $U(1)_D$  subgroup of the unbroken flavor symmetry,  $H$ , for the  $Sp(N_c)$ ,  $SO(N_c)$  and  $SU(N_c)$  gauge theories.

When the  $U(1)_D$  symmetry is gauged, the WZW-action includes additional gauge-interactions. Up to linear terms in the  $U(1)_D$  gauge boson  $\mathcal{A}$ , the gauged WZW-action is

$$\begin{aligned}\Gamma_{\text{WZW}}^{(1)} &= \frac{-i5N_c}{240\pi^2} \int_{M^4} e_D \mathcal{A} \text{Tr} Q [(d\Sigma\Sigma^{-1})^3 - (\Sigma^{-1}d\Sigma)^3] \\ &= \frac{-i2e_D N_c}{3\pi^2 f_\pi^3} \int d^4x \epsilon^{\mu\nu\rho\sigma} \mathcal{A}_\mu \text{tr} [\partial_\nu \Omega \partial_\rho \Omega \partial_\sigma \Omega + \partial_\nu \Omega \partial_\rho \mathcal{A} \partial_\sigma \mathcal{A}^\dagger].\end{aligned}\quad (4.4)$$

Here, ‘Tr’ in the first line is the trace over  $2N_f \times 2N_f$  matrices, while ‘tr’ in the second line is the trace over  $N_f \times N_f$  matrices. Semi-annihilation channels of  $\pi\pi \rightarrow \pi V$  are then present, and may spoil the SIMP mechanism. We suppress these processes by taking  $m_V \gtrsim 2m_\pi$ , rendering the process kinematically inaccessible during freeze-out (including effects of the thermal tail). Similarly, semi-annihilation interactions will be present from the gauged WZW action in the other gauge theories, and so we will always take  $m_V \gtrsim 2m_\pi$ .

For the  $SO(N_c)$  gauge theory with the  $SU(N_f)/SO(N_f)$  coset, if the number of flavors is even  $N_f = 2n_f$ , the gauged  $U(1)_D$  is embedded as

$$SU(n_f) \times U(1)_D \subset SO(N_f), \quad n_f = N_f/2 \quad (4.5)$$

The quarks transform as  $(\square, +1) \oplus (\bar{\square}, -1)$  under  $SU(n_f) \times U(1)_D$ , and the pions as  $(\square\square, +2) \oplus (\bar{\square}\bar{\square}, -2) \oplus (\text{adj}, 0)$ . This set of quantum numbers is consistent with the  $S$ -wave spin-zero bound states of massive Majorana quarks, with gauge indices contracted by the symmetric Kronecker-delta tensor, requiring the symmetric combination of flavor indices.

The generalization of the  $U(1)_D$  gauging in the  $SU(N_c)$  gauge theory in Ref. [8] is the following. The  $U(1)_D$  is embedded in the unbroken group as

$$SU(N_1) \times SU(N_2) \times U(1)_D \times U(1)_B \subset SU(N_f) \times U(1)_B \quad (4.6)$$

where the  $U(1)_D$  generator is

$$Q = \text{diag}(\underbrace{+1, \dots, +1}_{N_1}, \underbrace{-1, \dots, -1}_{N_2}) \quad (4.7)$$

with  $N_1 + N_2 = N_f$ . The pion states transform as  $(\square, \bar{\square}, 2) \oplus (\bar{\square}, \square, -2) \oplus (\text{adj}, 1, 0) \oplus (1, \text{adj}, 0) \oplus (1, 1, 0)$  under the  $SU(N_1) \times SU(N_2) \times U(1)_D$ .

The pions transforming in the non-trivial representations of the unbroken flavor group are protected from decay. The singlet pion in the  $SU(N_c)$  gauge theory, however, is not protected, but must be stable at the time of freeze-out; otherwise the other pions can scatter into this neutral state which then decays, eventually depleting the pion abundance. This issue cannot be avoided even if the singlet pion decays after freeze-out, because then the pion will decay after BBN and dissociate light elements. Therefore, the pion must be stable on cosmological time scales. The above choice of  $U(1)_D$ , Eq. (4.7), avoids the possibility of the singlet neutral pion  $\pi^0$  decaying into  $\pi^0 \rightarrow V^* V^* \rightarrow e^+ e^- e^+ e^-$  via the axial anomaly. This is because the singlet pion corresponds to the generator

$$T^0 = \sqrt{\frac{2}{(N_1 + N_2)N_1N_2}} \underbrace{(+N_2, \dots, +N_2)}_{N_1} \underbrace{(-N_1, \dots, -N_1)}_{N_2}, \quad (4.8)$$

and the triangle anomaly diagram vanishes,  $\text{Tr} T^0 Q Q = 0$ . The Adler–Bardeen theorem guarantees that this cancellation persists to all orders in perturbation theory [41]. Inserting more photons does not make  $\pi^0$  unstable. This is because charge conjugation invariance requires an even number of photons, while  $\text{Tr} T^0 Q^{2n} = 0$  for any  $n \in \mathbb{Z}$ . Double traces  $\text{Tr} T^0 Q^{2n-1} \text{Tr} Q^{2m-1}$  need not vanish, but the corresponding Feynman diagram, that may generate the effective operator, vanishes because of charged conjugation invariance of each fermion loop (Furry’s theorem). This is an intriguing case where a particle with no conserved quantum number appears to be stable. We note that the singlet  $\pi^0$  may decay through a higher dimension operator in the QCD-like theory, induced by some heavy particles, leading to a signal of late-time decay in the galactic halo.

The non-degeneracy among the charged and neutral pions, generated below the confinement scale, is of the order of  $\Delta m_\pi^2 \approx e_D^2 f_\pi^2$ . Sufficient degeneracy between the neutral and charged pions should be maintained in order to achieve the  $3 \rightarrow 2$  annihilation through the WZW term.

Having established the embedding of the dark photon in the confining SIMP theories, we now proceed to address the conditions of the SIMP mechanism in the presence of the dark photon.

## 4.2 SIMP conditions: Thermalization

A crucial requirement of the SIMP mechanism is that the entropy is dumped out of the SIMP sector while the SIMPs freeze-out through the  $3 \rightarrow 2$  process. We demand that the scattering rate between the SIMPs and the light Standard Model leptons ( $e^\pm, \nu, \bar{\nu}$ ) is faster than the expansion rate at the freeze-out temperature  $T_f$ , described by  $x_f = m_\pi/T_f \approx 20$ .

The scattering cross section  $\pi \ell \rightarrow \pi \ell$ , where  $\ell$  is a SM lepton (assumed massless), is

$$\begin{aligned} \sigma_{\text{scatt}}^f v_{\text{rel}} = & 8\pi\alpha_D\alpha \frac{\sum Q_\pi^2}{N_\pi} \frac{1}{c_\chi^2} \left[ - \left( \frac{c_\zeta^2}{m_V^2} + \frac{s_\zeta^2}{m_Z^2} \right) c_W t_\chi Q_f \right. \\ & \left. + \left( \frac{c_\zeta(s_\zeta + s_W t_\chi c_\zeta)}{m_V^2} - \frac{s_\zeta(c_\zeta - s_W t_\chi s_\zeta)}{m_Z^2} \right) \frac{1}{s_W c_W} (I_3^f - Q_f s_W^2) \right]^2 E_\ell^2 \quad (4.9) \end{aligned}$$

at threshold. We disagree with Eq. (18) in Ref. [8] where the mixing with the photon and  $Z$  are added incoherently in the cross section, or no  $Z$ -boson exchange diagram is considered. The above expression Eq. (4.9) should be summed over  $f = e_R^\pm, e_L^\pm, \nu_{e,\mu,\tau}, \bar{\nu}_{e,\mu,\tau}$ , where  $\langle E_\ell^2 n_\ell \rangle = \frac{12\zeta(5)}{\pi^2} \frac{15}{16} T_f^5$ . A naive estimate on the overall scattering rate can be obtained by requiring it be larger than the expansion rate at freezeout,

$$\Gamma_{\text{scatt}} = \sum_f \frac{\sigma_{\text{scatt}}^f v_{\text{rel}}}{E_f^2} \frac{12\zeta(5)}{\pi^2} \frac{15}{16} T_f^5 \lesssim H(T_f). \quad (4.10)$$

The resulting constraint is depicted by the dotted line of Fig. 2 (labeled **naive**).

A more detailed study of the rate of energy transfer between the SM and SIMP sectors results in a slightly more stringent constraint, strengthening the depicted curve [42]. The rate at which energy is lost for a dark matter particle scattering with a bath fermion via the process  $\pi_1 f_1 \rightarrow \pi_2 f_2$  is

$$\dot{E} = \frac{1}{2E_{\pi_1}} \int d\Pi_{f_1} d\Pi_{\pi_2} d\Pi_{f_2} (2\pi)^4 \delta^4(p_{\pi_1} + p_{f_1} - p_{\pi_2} - p_{f_2}) f_{f_1}(E_{f_1}) (E_{\pi_2} - E_{\pi_1}) |\mathcal{M}|^2, \quad (4.11)$$

where  $d\Pi_i = g_i d^3p_i / [(2\pi)^3 2E_i]$  and  $|\mathcal{M}|^2$  is averaged over initial and final state degrees of freedom. This should be compared to the rate at which kinetic-energy per-particle is being changed by the expansion of the universe, while chemical equilibrium is maintained by the  $3 \rightarrow 2$  self-annihilations. The rate at which the mass of the disappearing dark matter is transferred to kinetic energy in the dark-matter bath, by the  $3 \rightarrow 2$  process is,

$$\dot{E} = \frac{1}{n_\pi} \frac{\partial(m_\pi n_\pi)}{\partial t} \simeq -\frac{H m_\pi^2}{T}. \quad (4.12)$$

Therefore, in order for kinetic equilibrium to be maintained during freeze-out, the overall scattering rate must satisfy

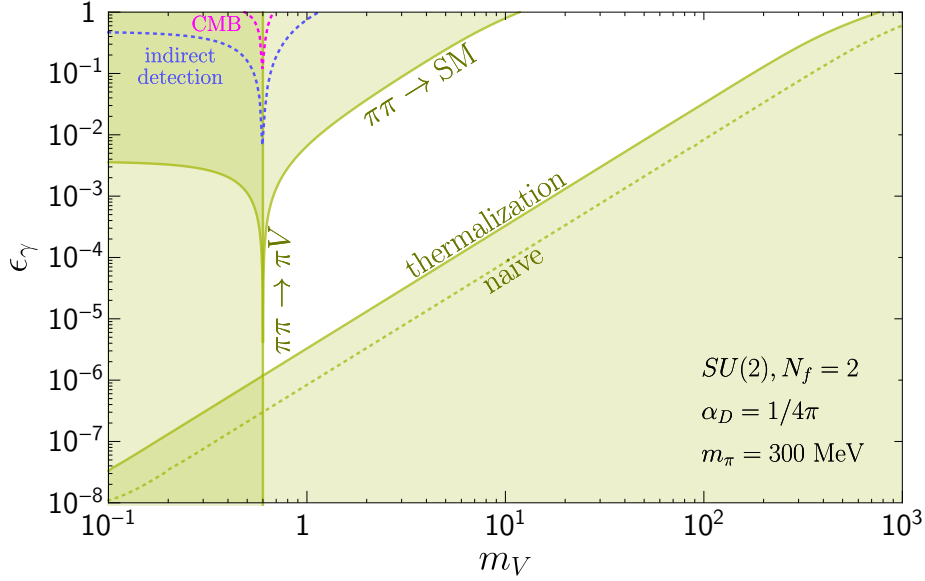
$$\frac{5\zeta(5)}{4} m_\pi \Gamma_{\text{scatt}} \lesssim \frac{H(T_f) m_\pi^2}{T_f}, \quad (4.13)$$

where the coefficient of the left-hand-side is determined by performing the integral in Eq. (4.11). The resulting constraint is depicted in the bottom shaded region of Fig. 2, stronger by a factor of  $\sim 3$  compared to the naive estimate.

### 4.3 SIMP conditions: Annihilation

The second condition for SIMP dark matter is that at the time of freeze out, no new SIMP annihilations beyond the  $3 \rightarrow 2$  process should be dominantly active. To this end, we require the  $2 \rightarrow 2$  annihilation of  $\pi\pi \rightarrow V^* \rightarrow e^+e^-$  to be out of equilibrium at freeze-out. The annihilation cross section is given by

$$\begin{aligned} \sigma_{\text{ann}}^\ell = & \beta \frac{8\pi\alpha_D\alpha}{3c_\chi^2} \frac{1}{N_\pi^2} \sum_\pi Q_\pi^2 m_\pi^2 \left[ - \left( \frac{c_\zeta^2}{m_V^2 - s} + \frac{s_\zeta^2}{m_Z^2 - s} \right) c_W t_\chi Q_\ell \right. \\ & \left. + \left( \frac{c_\zeta(s_\zeta + s_W t_\chi c_\zeta)}{m_V^2 - s} - \frac{s_\zeta(c_\zeta - s_W t_\chi s_\zeta)}{m_Z^2 - s} \right) \frac{1}{s_W c_W} (I_3^f - Q_\ell s_W^2) \right]^2, \end{aligned} \quad (4.14)$$



**Figure 2.** Constraints on the dark photon parameter space from requiring: (I) kinetic equilibrium between the Standard Model and SIMPs during freeze-out (labeled **thermalization**, and compared to the **naive** thermalization estimate); (II) dominance of the  $3 \rightarrow 2$  annihilation of SIMPs over  $\pi\pi \rightarrow V^* \rightarrow \ell^+\ell^-$  (labeled  $\pi\pi \rightarrow \text{SM}$ ), and (III) subdominance of the semi-annihilation process  $\pi\pi \rightarrow \pi V$  (labeled as such). We used the SIMPlest model of an  $SU(2) = Sp(2)$  gauge theory with  $N_f = 2$ ,  $\alpha_D = 1/4\pi$ , and  $m_\pi = 300$  MeV. (Suppressed) constraints from CMB data and indirect detection are shown in the dashed magenta and dashed blue curves, respectively.

and the requirement that annihilations are off during freezeout is

$$\Gamma_{\text{ann}} = \sum_{f=e_R^-, e_L^-, \nu_{e,\mu,\tau}} \frac{\sigma_{\text{ann}}^\ell}{\beta} 3x_f Y_\pi s_f \lesssim H(T_f). \quad (4.15)$$

Here,

$$s_f = \frac{2\pi^2}{45} g_{*s} T_f^3, \quad \text{and} \quad Y_\pi = \frac{n_\pi}{s} = 4.39 \times 10^{-10} \frac{\text{GeV}}{m_\pi}, \quad (4.16)$$

represent the entropy at freeze-out and the measured dark matter abundance, respectively. The resulting constraint is depicted in the top shaded region of Fig. 2.

The semi-annihilation process of  $\pi\pi \rightarrow \pi V$  must also be subdominant to the  $3 \rightarrow 2$  process during freeze-out. Since the thermalization requirement bounds  $\alpha_D \epsilon_\gamma^2$  from below, and the semi-annihilation process is proportional to  $\alpha_D$ , the only way to suppress the semi-annihilation process while maintaining thermal equilibrium with the SM is to require a kinematic suppression,  $Y_\pi Y_V \lesssim Y_\pi^3$  with the yields  $Y_{\pi,V} \sim e^{-m_{\pi,V}/T_f}$ . Therefore, we need  $m_V \gtrsim 2m_\pi$ , compensating for the thermal tail. This constraint is shown in the shaded region towards the left-hand side of Fig. 2.

In principle, stringent limits on light dark matter annihilations in the late Universe can exist [1]. These come from measurements of the CMB power spectrum [43] as well as from measurements of diffuse gamma-rays in the galactic halo [44]. In our case, the annihilation

of dark pions through the  $s$ -channel exchange of a dark photon,  $\pi^+\pi^- \rightarrow V^* \rightarrow e^+e^-$ , proceeds strictly in the  $p$ -wave, and the corresponding annihilation cross sections at later times are suppressed by  $v^2$ . Consequently, the cross section is much smaller at the time of recombination and in the galactic halo today. The process  $\pi^+\pi^- \rightarrow V^*V^* \rightarrow e^+e^-e^+e^-$  is also suppressed, by  $\epsilon_\gamma^4 (\frac{\alpha}{2\pi} \frac{m_\pi^2}{m_V^2})^2$ . The strength of the CMB and indirect-detection data is shown in the dashed magenta and blue and curves in Fig. 2, respectively. As is evident, these constraints play no role in the SIMP/dark-photon setup. Likewise, bounds from stellar cooling are evaded for SIMP pion masses of order a few hundred MeV, as they (and the heavier mediator) are too heavy to be produced in white dwarfs or supernovae. We learn that SIMP dark pions together with a dark photon mediator provides a light dark matter candidate that easily evades astrophysical and cosmological constraints.

## 5 Dark photon decays

The limits on the dark photon parameters discussed in Section 3 were obtained for 100% branching fraction either into visible or invisible modes. In a given model, the branching fractions into each mode must be computed in order to properly identify the viable parameter space. This is a complicated, yet fascinating task, due to the strong dynamics of both the QCD and SIMP sectors. We attempt the computation of partial widths in this section.

### 5.1 Invisible width of dark photon

The dark photon can decay into the dark sector. In  $e^+e^-$  annihilations, the virtual dark-photon can create hadronic states and probe the structure of the strongly coupled theory, determining properties such as the number of colors, flavors and the electric charges of the states. Given that the dark sector is a QCD-like strongly coupled gauge theory, the dark photon can produce vector resonances and hence perform spectroscopy of the dark sector, in a very similar way to spectroscopy of QCD resonances from virtual photons. To the best of our knowledge, this important point has not been discussed in the dark photon literature.

It is an interesting theoretical challenge how to model the dark spectroscopy. In what follows, we use a parametrization inspired by ‘soft wall’ holographic QCD [45] which mimics the resonance spectrum of radial excitations consistent with the Regge behavior of  $m_n^2 \propto (n + J)$ . This tool, used to model strongly coupled gauge theories, will enable us to understand what region of parameter space is available, as well as to demonstrate in a concrete way the various aspects of the proposed SIMP spectroscopy.

#### 5.1.1 Parton-level widths

When the dark photon mass is much higher than the dynamical scale of the hidden strong sector, the invisible decays of the hidden photon can be computed at parton-level into hidden quarks. The width of the hidden photon to quark-pairs is

$$\Gamma(V \rightarrow q\bar{q}) = \begin{cases} N_c N_f \Gamma_0 & SU(N_c) \text{ or } Sp(N_c), \\ \frac{1}{2} N_c N_f \Gamma_0 & SO(N_c), \end{cases} \quad (5.1)$$

where

$$\Gamma_0 = \frac{e_D^2}{12\pi} m_V = \frac{\alpha_D}{3} m_V, \quad (5.2)$$

is the two-body decay width into a single Dirac fermion. We note that the analysis of Ref. [8] considered vector decays into pions, rather than into quarks. In the  $SU(N_c = 3)$ ,  $N_f = 3$  model considered there, there are two charged scalars of charge two, and hence the invisible width to the charged pions is  $\Gamma(V \rightarrow \pi\pi) = 2\Gamma_0$ , while the invisible width into dark quarks is, according to Eq. (5.1), a factor of  $3N_c/2$  larger. Of course, the parton-level calculation breaks down near the dynamical scale where the dark quarks form dark hadrons, and will be treated in the next section.

Despite the factor of  $N_c N_f$  appearing in Eq. (5.1), we do not expect the invisible width to be very large. For instance, requiring that the  $U(1)_D$  coupling does not hit a Landau pole below the Planck scale gives

$$e_D^2 < \frac{8\pi^2}{b_0 \log(M_{\text{Pl}}/\Lambda)}, \quad (5.3)$$

where

$$b_0 = \begin{cases} \frac{4}{3} N_c N_f & SU(N_c) \text{ or } Sp(N_c), \\ \frac{2}{3} N_c N_f & SO(N_c). \end{cases} \quad (5.4)$$

Therefore,

$$\Gamma(V \rightarrow q\bar{q}) < \frac{\pi}{2 \log(M_{\text{Pl}}/\Lambda)} m_V \simeq 0.037 m_V. \quad (5.5)$$

The perturbativity constraint on the  $U(1)_D$  coupling, Eq. (5.3), also ensures that the dark pions are sufficiently degenerate for the  $3 \rightarrow 2$  (co-annihilation) process to proceed.

### 5.1.2 Vector resonances

As the mass of the hidden photon approaches the confinement scale, the effects of hadronic vector resonances of the strongly coupled theory on the invisible decay width become important. However, we cannot compute the strongly coupled physics from first principles. For the purpose of calculating the invisible branching ratios of the hidden photon, and illustrating spectroscopy at future  $e^+e^-$  colliders, the resonance structure needs to be modeled in a reasonable way. For the discussions here, we focus on  $Sp(N_c)$  gauge theories, while  $SO(N_c)$  and  $SU(N_c)$  gauge theories are discussed in Appendix A.

The  $\rho$  mesons are  $L = 0$ ,  $S = 1$  bound states of the Majorana quarks whose gauge indices are contracted by the (anti-symmetric) symplectic tensor. Therefore the flavor indices transform as a symmetric tensor of the unbroken  $Sp(2N_f)$ . Under  $SU(N_f) \times U(1)_D \subset Sp(2N_f)$ , they decompose as  $(\square\square, +2) \oplus (\overline{\square}\overline{\square}, -2) \oplus (\text{adj}, 0) \oplus (1, 0)$ . The dark photon couples to the singlet representation  $\rho(1, 0)$ .

The width of the dark photon can be obtained from the vector-vector correlation function or vacuum polarization diagram,

$$m_V \Gamma_V = N_f e_D^2 m_V^2 \Im m \Pi(m_V^2), \quad (5.6)$$

where the factor of  $N_f$  comes from the normalization of the  $U(1)_D$  charge, and the vacuum polarization function  $\Pi(q^2)$  is defined by

$$\int d^4x e^{iq \cdot x} \langle J_\mu^a(x) J_\nu^b(0) \rangle = i \delta^{ab} (q_\mu q_\nu - g_{\mu\nu}) \Pi(q^2). \quad (5.7)$$

In the strongly coupled regime and large  $N_c$ , the same vacuum polarization function should be obtained by summing over the vector resonances [46]. The vector resonances  $\rho_n$  are characterized by their masses and decay constants with the corresponding vector currents  $J_\mu^a$ ,

$$\langle 0 | J_\mu^a | \rho_n^b \rangle = \delta^{ab} F_{\rho_n} \epsilon_\mu. \quad (5.8)$$

Casting this into an effective Lagrangian using its source gauge field  $J_\mu^a \mathcal{A}^{\mu a}$ ,

$$\mathcal{L}_{eff} = -\frac{1}{4e_D^2} \mathcal{A}_{\mu\nu}^a \mathcal{A}^{\mu\nu a} + \sum_n -\frac{1}{4} \rho_{\mu\nu}^a \rho^{\mu\nu a} + \frac{F_{\rho_n}}{2m_{\rho_n}^2} \mathcal{A}_{\mu\nu}^a \rho_n^{\mu\nu a} + \frac{1}{2} m_{\rho_n}^2 \rho_n^a \rho_n^{\mu a}. \quad (5.9)$$

The vacuum polarization diagram is obtained by integrating out  $\rho_n$  and differentiating the resulting action with respect to  $\mathcal{A}_\mu^a$  twice (right-most diagram in Fig. 8 in Appendix A),

$$\begin{aligned} (q_\mu q_\nu - g_{\mu\nu}) \Pi(q^2) &= \sum_n \left( \frac{F_{\rho_n}}{m_{\rho_n}^2} \right)^2 (q_\mu q_\nu - g_{\mu\nu} q^2) \frac{-g^{\kappa\lambda} + q^\kappa q^\lambda / m_{\rho_n}^2}{q^2 - m_{\rho_n}^2 + i\Gamma_n^\rho m_{\rho_n}} (q_\lambda q_\nu - g_{\lambda\nu} q^2) \\ &= -\sum_n \left( \frac{F_{\rho_n}}{m_{\rho_n}^2} \right)^2 (q_\mu q_\nu - g_{\mu\nu} q^2) \frac{q^2}{q^2 - m_{\rho_n}^2 + i\Gamma_n^\rho m_{\rho_n}}, \end{aligned} \quad (5.10)$$

leading to

$$\Pi(q^2) = -\sum_n \frac{F_{\rho_n}^2}{m_{\rho_n}^2} \frac{q^2}{(q^2 - m_{\rho_n}^2 + i\Gamma_n^\rho m_{\rho_n}) m_{\rho_n}^2}. \quad (5.11)$$

In order to perform the sum, the spectrum and decay constants of the  $\rho$ -mesons are needed. To calculate these, we borrow results from the ‘soft-wall’ holographic QCD [45], the minimal details of which are reviewed in Appendix A, and state the main results here. Holographic QCD identifies  $\rho$ -mesons as the gauge bosons of the gauge flavor symmetry in the AdS bulk, which precisely agrees with the quantum numbers discussed above. The spectrum and decay constants are given by Eqs. (A.4) and (A.8) (in units of AdS curvature),

$$m_{\rho_n}^2 = 4n, \quad \frac{F_{\rho_n}^2}{m_{\rho_n}^2} = \frac{2}{g_5^2}, \quad g_5^2 = \frac{12\pi^2}{N_c}, \quad n = 1, 2, \dots \quad (5.12)$$

Performing the sum over the tower of  $\rho$ -meson resonances in Eq. (5.11), gives

$$\Pi(q^2) - \Pi(0) = -\frac{N_c}{24\pi^2} \left( 1 - \frac{1}{\pi} \tan^{-1}(N_f - 1) \frac{g_5^2}{96\pi} \right)^{-1} H_{-q_{eff}^2/4}, \quad (5.13)$$

with  $H$  the harmonic number [see Eq. (A.9)] and

$$q_{eff}^2 = q^2 \left( 1 + i \frac{\Gamma_n^\rho}{m_{\rho_n}} (m_\rho^2 = q^2) \right). \quad (5.14)$$

With simplifying assumptions given in Eq. (A.11), the  $\rho \rightarrow \pi\pi$  partial widths are given by

$$\Gamma_n^\rho = (N_f - 1) \frac{\beta^3}{96\pi} g_5^2 m_{\rho_n}, \quad \beta = \sqrt{1 - \frac{4m_\pi^2}{m_{\rho_n}^2}}. \quad (5.15)$$



## 5.2 Visible widths of dark photon

The decay of the dark photon into the visible sector can be computed reliably (see *e.g.* Ref. [47]). For the leptonic final states:

$$\Gamma(V \rightarrow \ell^+ \ell^-) = \frac{\alpha \epsilon_\gamma^2}{\alpha_D} \frac{\beta(3 - \beta^2)}{2} \Gamma_0, \quad \beta = \sqrt{1 - \frac{4m_\ell^2}{m_V^2}} \quad (5.16)$$

where  $\Gamma_0 \equiv \alpha_D m_V / 3$  as in Eq. (5.2). The hadronic width is modulated by the resonance structure of QCD. Using the standard notation of the hadronic cross section in  $e^+ e^-$  annihilations at center of mass energy  $\sqrt{s}$ ,

$$R(s) \equiv \frac{\sigma(e^+ e^- \rightarrow \text{hadron})}{\sigma(e^+ e^- \rightarrow \mu^+ \mu^-)_0}, \quad (5.17)$$

where the subscript 0 refers to the lowest order QED calculation for massless muons,

$$\sigma(e^+ e^- \rightarrow \mu^+ \mu^-)_0 = \frac{4\pi\alpha^2}{3s}, \quad (5.18)$$

the hadronic width is

$$\Gamma(V \rightarrow \text{hadrons}) = \frac{\alpha \epsilon_\gamma^2}{\alpha_D} R(m_V^2) \Gamma_0. \quad (5.19)$$

We use data of  $R(s)$  from Ref. [48].

When  $m_V$  is much higher than the QCD scale and is comparable to or above  $m_Z$ , the full expression of Eq. (3.5) should be used for each Weyl fermion in the Standard Model,

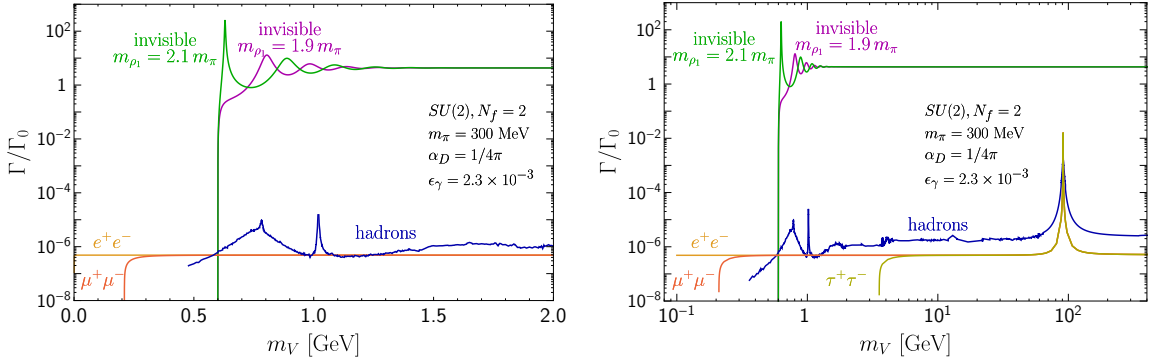
$$\Gamma_{V \rightarrow f \bar{f}} = \frac{\alpha}{\alpha_D} \frac{1}{2} \left[ Q_f \epsilon_\gamma + \frac{1}{s_W c_W} (I_3^f - Q_f s_W^2) \epsilon_Z \right]^2 \frac{\beta(3 - \beta^2)}{2} \Gamma_0. \quad (5.20)$$

For quark final states, this should be multiplied by the QCD correction factor,  $1 + \alpha_s/\pi$ .

## 5.3 Results and constraints

We are now ready to compute the partial widths of the dark photon. If the dark photon is heavier than multi-GeV, it is in the perturbative regime for both our QCD and the QCD-like dark-sector. For a light(er) dark photon, non-perturbative effects are important and the partial widths are highly non-trivial.

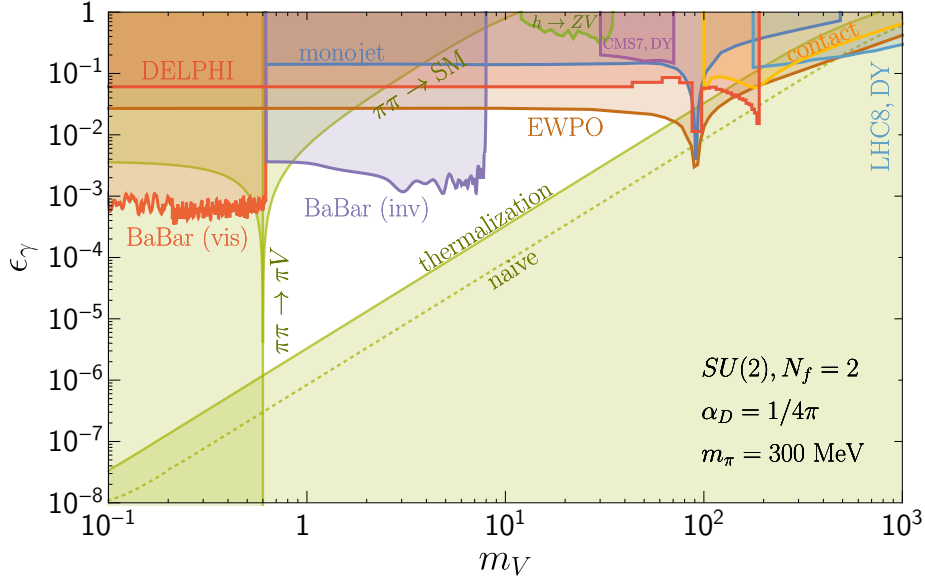
The last ingredient needed is the mass scale of the  $\rho$  resonances relative to  $m_\pi$ ; this is highly model-dependent. In the SIMplest model of an  $Sp(2) \simeq SU(2)$  gauge group with  $N_f = 2$ , perturbativity considerations along with self-scattering constraints point to the rough scales of  $m_\pi/f_\pi \approx 6$  and  $m_\pi \approx 300$  MeV [3]. Since the theory is in the very strongly interacting regime, this is understood as a proxy to the scales involved;  $\mathcal{O}(1)$  corrections are likely to apply. In what follows, we use these estimates for the pion scales involved for illustration purposes. We do not know what the corresponding  $\rho$ -meson mass is, however we do expect the first of the vector-meson states  $\rho_n$  to at least be heavier than  $\pi$ . This is because, in the non-relativistic quark model, the mass splitting between the mesons is due to the hyperfine interaction of two quarks inside the meson, which makes the  $F = 1$



**Figure 3.** Partial widths of the dark photon into various final states, as a function of dark photon mass: invisible (purple and green), hadrons (blue), electrons (orange), muons (red) and taus (dark yellow, right panel only). We use the SIMPlEst model of an  $SU(2)$  gauge group with  $N_f = 2$  (four doublets), with  $m_\pi = 300$  MeV,  $\alpha_D = 1/(4\pi)$  and  $\epsilon_\gamma = 2.3 \times 10^{-3}$ . The partial width on the vertical axis is normalized to  $\Gamma_0 = \alpha_D m_V/3$ . For the dark sector, we present two benchmark cases for  $\rho$  masses:  $m_{\rho_1} = 1.9m_\pi\sqrt{n}$  MeV (solid purple) and  $2.1m_\pi\sqrt{n}$  (solid green), to illustrate the behavior of different possible dynamics. In the former case, the first state  $\rho_1$  does not show up in the invisible width because it is lighter than  $2m_\pi$ ; this is not the situation for the latter case. The right panel simply extends the mass range of the hidden photon to heavier masses than shown in the left panel.

state higher than the  $F = 0$  state. The mass splitting narrows with increased quark mass. This is indeed the case in QCD when going from  $\rho$  and  $\pi$ ,  $\phi$  and  $\eta$  to  $J/\psi$  and  $\eta_c$ . The vector meson is always heavier than the pseudoscalar. As a result, we will always take the  $\rho$ 's to be heavier than the pions. Moreover, suppressing semi-annihilation processes of  $\pi\pi \rightarrow \pi\rho$  (and in the  $SU(N_c)$  case, also  $\pi\pi \rightarrow \pi\omega$ ) compared to the  $3 \rightarrow 2$  process requires  $m_\rho \gtrsim 2m_\pi$ .

In Fig. 3 we plot the various branching fractions of the dark photon as a function of its mass. The left and right panels of Fig. 3 correspond to the low-mass and high-mass ranges of the dark photon, respectively. Here we take  $m_\pi = 300$  MeV,  $\alpha_D = 1/(4\pi)$  and  $\epsilon_\gamma = 2.3 \times 10^{-3}$ , corresponding to the largest allowed value of  $\epsilon_\gamma$  for the low mass range, as will be seen below in Fig. 4. We choose two representative benchmark cases for lightest  $\rho$ -meson mass:  $m_{\rho_1} = 1.9m_\pi$  with the corresponding tower, and  $m_{\rho_1} = 2.1m_\pi$  with the corresponding tower. Given that the  $m_\rho/m_\pi$  mass ratio cannot be computed reliably, these two benchmark cases are sufficient as examples of different potential features. As mentioned above, we do not consider  $\rho$  masses much below  $2m_\pi$ , such that semi-annihilations into  $\rho$ 's do not dominate the relic abundance, and are at most comparable to the  $3 \rightarrow 2$  rate. If the mass of some low-lying vector state does not exceed  $2m_\pi$ , this resonance cannot decay invisibly and instead decays back to SM particles. This is exemplified by the  $m_{\rho_1} = 1.9m_\pi$  case, plotted in purple, where the first low-lying state decays into SM modes. Alternatively, the lightest low-lying vector state can be heavier than two pions, in which case all invisible modes are available. Such behavior is represented by the  $m_{\rho_1} = 2.1m_\pi$  case, which is plotted in green.



**Figure 4.** Combined requirements and constraints for SIMP dark matter with a dark photon mediator. The results are plotted for the SIMplest model of an  $SU(2)$  gauge group with  $N_f = 2$ ,  $m_\pi = 300$  MeV and  $\alpha_D = 1/(4\pi)$ . The shaded green regions define the range of validity of the SIMP mechanism with a dark photon mediator to the visible sector, as in Fig. 2. The other colored shaded regions are the experimental constraints on  $\epsilon_\gamma$ , coming from visible [33] and invisible [38] decays at BaBar, mono-photon searches at DELPHI [40], electroweak precision observables [29], Drell-Yan production measured at ATLAS [36] and CMS [35], four-fermion contact operators at LEP II [31],  $h \rightarrow ZV \rightarrow 4\ell$  from CMS data [30], and monojet searches at CDF and ATLAS [39].

Having computed the partial widths, the dark photon constraints of Fig. 1 can be appropriately weighted and combined with the SIMP constraints of Fig. 2 to obtain the viable parameter space. The results for the SIMplest model of an  $Sp(2) = SU(2)$  gauge group with  $N_f = 2$  are shown in Fig. 4. The unshaded region in the center is phenomenologically viable. For instance, a dark photon with mass of a few GeV can have  $\epsilon_\gamma$  as large as  $\sim 10^{-3}$  but not smaller than  $\sim 10^{-5}$ ; the  $\epsilon_\gamma$  range mostly broadens (shrinks) for lighter (heavier) masses. Concerning the BaBar limits, our treatment is as follows. Given the smallness of the decay widths into SM particles compared to the decay width into dark resonances (see Fig. 3), we take 100% branching fraction into SM particles when  $m_V < 2m_\pi$ , and once decays into the hidden sector are kinematically allowed, we use 100% branching fraction into the hidden sector. The depicted results use  $m_\pi = 300$  MeV and  $\alpha_D = 1/(4\pi)$ , and the effect of altering the dark matter mass can be readily understood. As  $m_\pi$  varies, the constraint from suppressing semi-annihilation  $m_V \gtrsim 2m_\pi$  shifts accordingly, as does the resonant annihilation dip near  $m_V = 2m_\pi$ . Likewise, the transition point between the BaBar SM final state searches and invisible modes will shift accordingly. Qualitatively, the viable parameter space of SIMPs shown in Fig. 4 well-captures the viable parameter space of other gauge groups with different number of colors and flavors as well.

## 6 SIMP spectroscopy

If the dark-photon is produced off-shell, then its production probes the spectrum of resonances of the SIMP dynamics. The idea presented here applies more broadly to any strongly coupled sector that interacts in a similar way with the SM.

There are two distinct regimes to the SIMP physics that depend on whether the singlet dark  $\rho$ -meson is kinematically allowed to decay into the dark pions, or decays via the hidden-photon into SM states. All but the singlet  $\rho$ -mesons transform nontrivially under the broken flavor symmetry (see Table 2), and may be stable against all decays, in the  $Sp(N_c)$  and  $SO(N_c)$  theories. For the  $SU(N_c)$  case, since the  $\rho$ -mesons and pions transform identically, the  $\rho$ 's can decay into  $\pi + V^{(*)} \rightarrow \pi + \ell^+ \ell^-$  etc.

When the dark  $\rho$ -meson can decay into dark pions, then any production of dark quarks or dark hadrons will quickly cascade decay into the dark pions, resulting in the dynamics of the dark sector being primarily measured indirectly. Since the SIMP parameter space naturally pushes  $m_q/f_\pi > 1$  [3], and therefore  $m_\pi \rightarrow m_\rho$ , it is quite possible that the dark  $\rho$ -mesons will be stable against strong decays inside the dark sector.

If the dark  $\rho$ -meson is kinematically forbidden to decay into the dark pions, then dark  $\rho$ -mesons must decay via interactions with the dark photon. The singlet  $\rho$ -mesons can decay via  $\rho \rightarrow V^{(*)} \rightarrow \text{SM}$ . This can lead to a lepton-jet-like signal [49].

There are additional low-lying hadronic states that may play a role in the phenomenology. The most important one is the  $a_1$  resonance,<sup>2</sup> which is the  $L = 1$ ,  $S = 1$  spin-1 resonance. The  $a_1$  transforms under the flavor symmetries as the pions. When the  $a_1$  state is kinematically allowed to decay into the strong sector, it may decay via the dominant processes  $a_1 \rightarrow 3\pi, \pi\rho_1$ . If the lowest  $\rho$ -mesons strong decays are forbidden, then strong  $a_1$  decays will be forbidden as well. Instead the  $a_1$  will be forced to decay via  $a_1 \rightarrow \pi V^{(*)}$ . This decay is always allowed, and so the  $a_1$  will not be stable. Similar statements can be made about the  $\omega$  resonances in  $SU(N_c)$  theories.

Additional states like the  $L = 1$ ,  $S = 0$  states ( $b_1$  meson) are stable to decays to dark pions in  $Sp(N_c)$  or  $SO(N_c)$  theories. They will be able to decay with a dark photon into dark  $\rho$ -mesons. The number of states that are stable to strong decays grows as dark quark masses grow and the theory transitions into a quirk-like regime [50].

### 6.1 Measuring invisible mass

In  $e^+e^- \rightarrow \gamma + \text{inv}$  annihilations, the energy of the recoil photon is in one-to-one correspondence with the invariant mass  $M_{\text{inv}}$  of the invisible system,

$$E_\gamma = \frac{\sqrt{s}}{2} \left( 1 - \frac{M_{\text{inv}}^2}{s} \right). \quad (6.1)$$

---

<sup>2</sup>Note that the subscript here refers to  $J = 1$ , not the radial excitations  $(a_1)_n$ . The same comment applies to  $b_1$ .

The production cross section is given by

$$\sigma(e^+e^- \rightarrow \gamma + \text{inv}) = \frac{3\alpha}{s} \int d\cos\theta \int dM_{\text{inv}}^2 \frac{M_{\text{inv}}^2}{(M_{\text{inv}}^2 - m_V^2)^2 + m_V^2 \Gamma^2} \times \frac{\Gamma(V \rightarrow \text{inv})}{M_{\text{inv}}} \frac{\Gamma(V \rightarrow e^+e^-)}{M_{\text{inv}}} \frac{8 - 8\beta + 3\beta^2 + \beta^2 \cos 2\theta}{\beta \sin^2 \theta}, \quad (6.2)$$

with  $\beta = 1 - M_{\text{inv}}^2/s$ , which agrees with the result of Ref. [51]. Here, the widths are to be computed as if  $m_V = M_{\text{inv}}$ , reflecting the off-shell nature of the dark photon in the process, using the  $m_V$  dependence of widths given in Fig. 3. We take the acceptance of photons to be  $|\cos\theta| < \cos 12^\circ$ . This is only slightly optimistic compared to the actual geometric coverage of 12 to 157 degrees at Belle-II. Belle-II is working on implementing a single photon trigger [52], a significant effort considering the high rate of bremsstrahlung background.

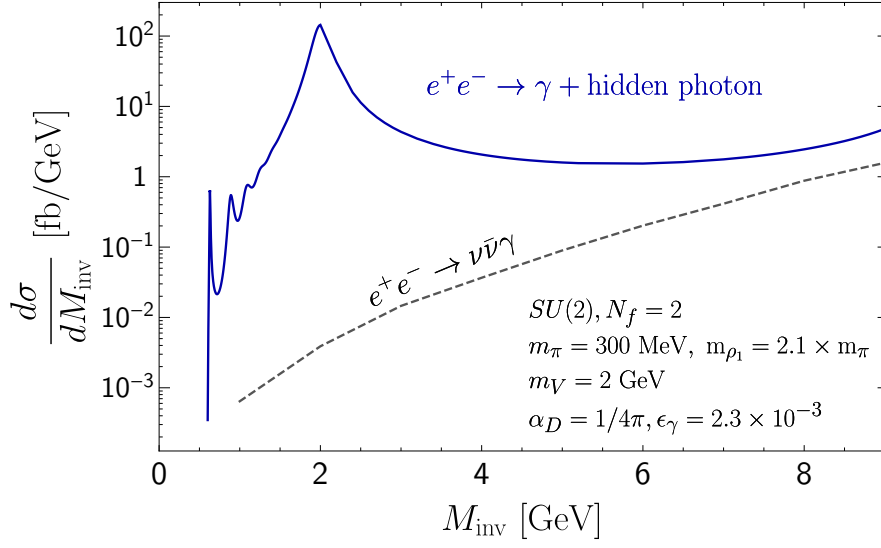
Using the partial widths modeled in Section 5, we show the invariant-mass spectrum in  $e^+e^-$  collisions at  $\sqrt{s} = 10$  GeV. The entire mass range is shown in Fig. 5, while Fig. 6 focuses on the low mass range. The dark photon peak in the invariant mass distribution is clearly visible (here at 2 GeV). The irreducible background of  $e^+e^- \rightarrow \gamma\nu\bar{\nu}$  is shown in the dashed gray curve in Fig. 5. For the low masses in Fig. 6, we present two possible dynamics, using the benchmarks of  $m_{\rho_1} = 1.9m_\pi$  and  $2.1m_\pi$  as representative cases. In the latter case, when all  $\rho$  resonances are above the two-pion threshold, their decays are entirely into the dark sector (green curve). In the former case, only the second and higher vector resonances decay invisibly (purple curve), while the first vector resonance cannot decay into dark pions, and so decays into charged leptons and QCD pions. We describe these visible decays next.

If the dark photon mass is well above the confinement scale, it decays into dark quarks which fragment dominantly into dark mesons. As discussed above, many of the meson states either decay into dark pions, are stable against the strong decays, or are completely stable. However, a fraction of the produced mesons can decay back into the Standard Model and create additional lepton pairs or QCD-hadrons. The study of the invisible mass spectrum therefore needs to be done paying careful attention to this possibility that a fraction of the ‘invisible mass’ is actually visible.

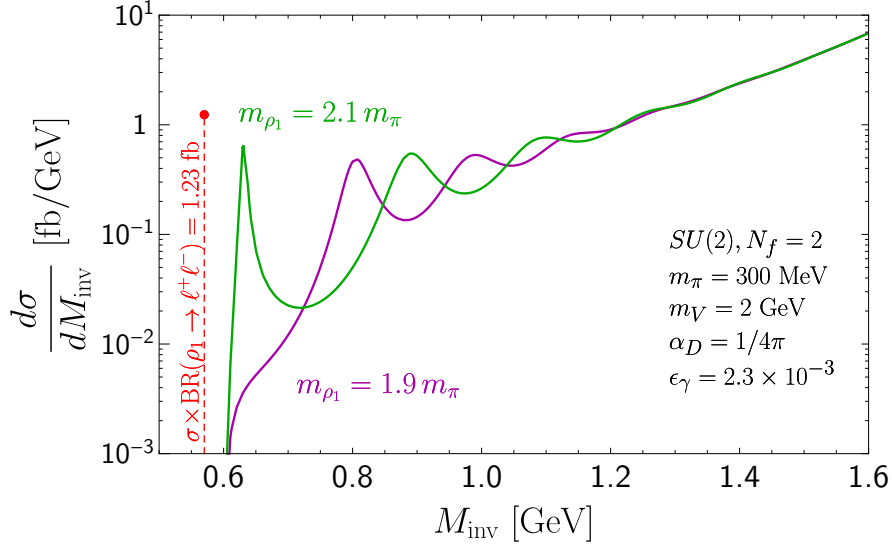
Finally, various dark sector masses can be measured without full knowledge of the strongly coupled sector. The invariant mass of lepton pairs would measure the mass splittings among the SIMP resonances, as well as the lowest  $\rho$  states. The cutoff at small invariant mass in the  $M_{\text{inv}}$ -distribution corresponds to  $2m_\pi$ , measuring the mass of the dark matter. A combination of these techniques can in principle reveal the spectrum of the SIMP sector resonances.

## 6.2 Visible widths of SIMP resonances

If a resonance in the dark sector cannot decay ‘strongly’, say when  $m_\rho < 2m_\pi$ , it will decay into the visible sector. This is akin to the fact that  $J/\psi$  cannot decay into  $D^+D^-$ , and decays instead into  $\mu^+\mu^-$  with a large branching fraction ( $\sim 6\%$ ). In the case of  $SU(N_c)$



**Figure 5.** The cross section for  $e^+e^- \rightarrow \gamma + \text{dark photon}$  at  $\sqrt{s} = 10\text{GeV}$ , where the dark photon decays invisibly (solid blue curve). We plot this cross section in fb/GeV as a function of the invariant mass of the invisible system. The dashed gray line is the irreducible background coming from  $e^+e^- \rightarrow \gamma + Z^*$  with  $Z^* \rightarrow \nu\bar{\nu}$ . Here we take the SIMplest model of an  $SU(2)$  gauge group and  $N_f = 2$  (four doublets), with  $m_\pi = 300$  MeV,  $m_{\rho_1} = 2.1 \times m_\pi$ ,  $\alpha_D = 1/(4\pi)$  and  $\epsilon_\gamma = 2.3 \times 10^{-3}$ . We use a dark photon mass of 2 GeV, and its peak is clearly visible in the  $M_{\text{inv}}$  spectrum.



**Figure 6.** The same as in Fig. 5, but focusing on the low invariant mass region. Two benchmarks are considered. When  $m_{\rho_1} = 2.1m_\pi$  (solid green), all  $\rho$  states decay into  $\pi\pi$ . When  $m_{\rho_1} = 1.9m_\pi$  (solid purple), the first  $\rho$  state cannot decay invisibly and decays instead into charged leptons. These decays are described in the dashed red vertical line, and the corresponding  $\sigma \times \text{BR}$  is marked as well.

gauge theories,  $\omega$  cannot decay into two pions but can into three pions. Therefore the  $\omega$  states are more likely to decay into the Standard Model leptons.

The leptonic branching fraction for a dark resonance can be computed in the same way as for the vector mesons of ordinary QCD. For the ordinary  $\rho$ -meson of the SM,

$$\begin{aligned}\langle 0|J_{EM}^\mu|\rho_{\text{SM}}\rangle &= \langle 0|e\left(\frac{2}{3}\bar{u}\gamma^\mu u - \frac{1}{3}\bar{d}\gamma^\mu d\right)|\rho_{\text{SM}}\rangle = e\langle 0|\frac{1}{2}(\bar{u}\gamma^\mu u - \bar{d}\gamma^\mu d)|\rho_{\text{SM}}\rangle \\ &= eF_{\rho_{\text{SM}}}\epsilon^\mu.\end{aligned}\quad (6.3)$$

As a result, for the SM  $\rho$ -meson

$$\Gamma(\rho_{\text{SM}} \rightarrow e^+e^-) = \frac{4\pi}{3} \frac{\alpha^2 F_{\rho_{\text{SM}}}^2}{m_{\rho_{\text{SM}}}^3}. \quad (6.4)$$

Following the SM example, but for the case of  $Sp(N_c)$  theories and the singlet  $\rho$ -mesons (the vector corresponding to the gauge  $U(1)_D$ -symmetry, up to a normalization),

$$\langle 0|J_D^\mu|\rho_n\rangle = \langle 0|e_D \sum_{i=1}^{N_f} \bar{q}_i \gamma^\mu q_i|\rho_n\rangle = e_D \sqrt{2N_f} F_{\rho_n} \epsilon^\mu. \quad (6.5)$$

Here,  $q_i$  are regarded as Dirac fermions. We take the decay constants from the soft wall model [45],

$$F_{\rho_n}^2 = \frac{m_{\rho_n}^2 m_{\rho_1}^2 N_c}{24\pi^2}, \quad (6.6)$$

see Appendix A for more details. Taking into account the mixing between the dark photon and our photon, and considering the  $SU(N_c)$  and  $SO(N_c)$  gauge theories as well, the widths into dilepton final states are

$$\Gamma(\rho_n \rightarrow \ell^+ \ell^-) = \frac{\alpha_D \epsilon_\gamma^2 \alpha}{18\pi} \frac{m_{\rho_1}^2}{m_{\rho_n}} \frac{\beta(3 - \beta^2)}{2} N_c N_f^{\text{eff}}, \quad (6.7)$$

where

$$N_f^{\text{eff}} = \begin{cases} 2N_f & Sp(N_c) \\ N_f & SO(N_c) \\ \frac{8N_1 N_2}{N_1 + N_2} & SU(N_c) \end{cases}. \quad (6.8)$$

In the case of  $SU(N_f)$  theories, there is also the  $\omega_n$ , which is degenerate with the  $\rho_n$  in holographic QCD, with width

$$\Gamma(\omega_n \rightarrow e^+e^-) = \frac{\alpha_D \epsilon_\gamma^2 \alpha}{18\pi} \frac{m_{\rho_1}^2}{m_{\rho_n}} \frac{\beta(3 - \beta^2)}{2} N_c \frac{2(N_1 - N_2)^2}{N_f}. \quad (6.9)$$

Finally, there are widths into (our) hadrons, given by the multiplication of the above expressions by  $R(s)$  of Eq. (5.17) with  $\beta = 1$ .

We can use the production cross section in Eq. (6.2) by identifying the vector  $V$  not as the dark photon but rather as SIMP resonances, in the narrow width approximation:

$$\int dM^2 \frac{1}{(M^2 - m^2)^2 + m^2 \Gamma^2} = \frac{\pi}{m\Gamma}. \quad (6.10)$$

Then Eq. (6.2) reduces to

$$\begin{aligned} & \frac{d\sigma(e^+e^- \rightarrow \gamma\rho \rightarrow \gamma\ell^+\ell^-)}{d\cos\theta} \\ &= \frac{\alpha^2\epsilon_\gamma^2\alpha_D}{4s} \frac{m_{\rho_1}^2}{m_{\rho_n}^2} \frac{8 - 8\bar{\beta} + 3\bar{\beta}^2 + \bar{\beta}^2\cos 2\theta}{\bar{\beta}\sin^2\theta} \text{BR}(\rho_n \rightarrow \ell^+\ell^-) N_c N_f^{\text{eff}}, \end{aligned} \quad (6.11)$$

with  $\bar{\beta} = 1 - m_{\rho_n}^2/s$ . The results are shown in Fig. 6 in the red vertical line, for the first vector resonance of the  $m_{\rho_n} = 1.9m_\pi n^{1/2}$  tower for the  $SU(2)$  gauge theory with  $N_f = 2$ .

## 7 Future prospects

We now discuss future prospects for SIMP dark matter searches interacting with the SM via a kinetically mixed hidden photon. These include direct detection, spectroscopy at  $e^+e^-$  colliders, hadron collider signatures, ILC prospects and beam dump experiments. To illustrate the many future probes of the setup, in Fig. 7 we show projections of various future measurements into the viable parameter space of Fig. 4. The shaded green and gray regions correspond to the excluded regions of Figs. 2 and 4, accordingly, while the colored curves show future reach, as detailed below.

### 7.1 Dark matter direct detection

Sub-GeV dark matter is challenging for direct detection experiments relying on conventional nuclear recoil signals (though see Refs. [53–55]). For light dark matter of order hundreds of MeV, scattering off an electron will typically deposit  $m_e v_{\text{halo}}^2 \simeq 100$  eV energy, and can excite electrons in a semiconductor above the band gap, leading to a detectable signal [56, 57]. (In this context, see Ref. [58, 59] for use of superconducting targets as well.)

The non-relativistic scattering cross section of SIMP pions on electrons is given by

$$\sigma = \frac{1}{8\pi} \left( \frac{2\epsilon_\gamma e e_D}{m_V^2} \right)^2 s \bar{\beta}^2, \quad \bar{\beta} = 1 - \frac{m_\pi^2}{s}. \quad (7.1)$$

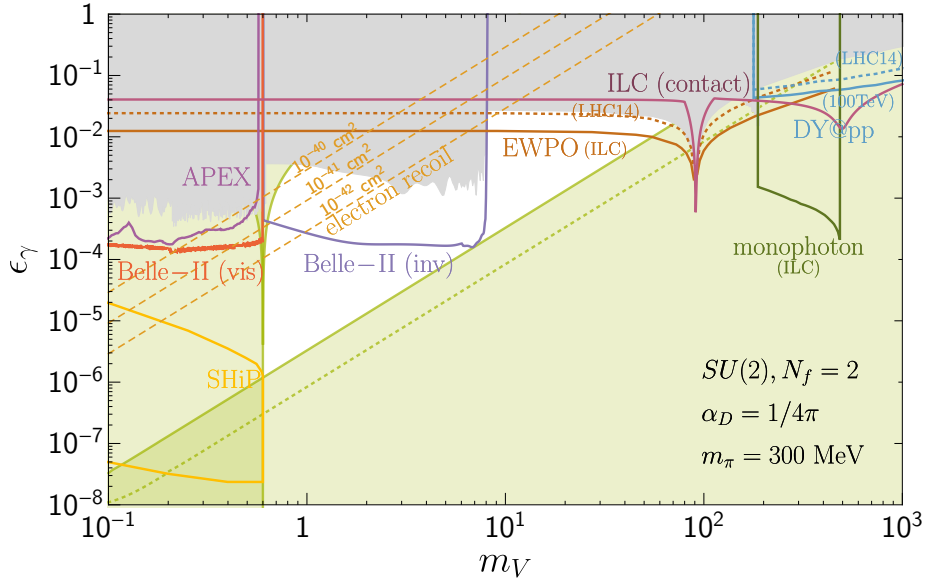
Here,  $s = (m_e + m_\pi)^2 + m_e m_\pi v_\pi^2$ .

For dark matter mass of order a few hundred MeV, the current direct detection constraint is  $\sigma_{\text{DD}} \lesssim 10^{-35} \text{ cm}^2$  [60] using data from the Xenon10 experiment [61]. The analysis of Ref. [62] suggests that an improved reach down to cross sections of  $\mathcal{O}(10^{-41} \text{ cm}^2)$  may be achievable in a silicon or germanium semiconductor detector with a kg-year of exposure. Contours of the SIMP-electron elastic scattering cross section are overlaid in Fig. 7 in dashed orange curves.

### 7.2 Spectroscopy at low-energy lepton colliders

At SuperKEKB, an integrated luminosity of  $50 \text{ ab}^{-1}$  is anticipated. In the Belle II experiment, the implementation of a single photon trigger has started to be studied in order to make the beam-induced and QED-originated backgrounds under control. The expected energy resolution of the electromagnetic calorimeter is 1.5% at  $E_\gamma = 5 \text{ GeV}$ ,





**Figure 7.** Anticipated future reach of various experiments into the viable parameter space of the SIMP/dark photon setup. These are: Belle-II visible [65] (solid red) and invisible [29] (solid purple); SHiP [66] (solid orange); Drell-Yan at LHC14 [30] (dashed blue) and 100 TeV  $pp$  collider [30] (solid blue); ILC monophoton search [67] (solid green), reinterpretation of contact operators at the ILC [68, 69] (solid violet) and EWPO at the LHC (dashed brown) and the ILC [29, 30] (solid brown); direct detection via electron recoil signals (dashed orange); and APEX [70] (solid dark purple).

2% at  $E_\gamma = 1$  GeV and 4% for  $E_\gamma = 0.1$  GeV [52] (see also Ref. [63]). There is a large background at  $M_{\text{inv}} = 0$  when one of the photons in the  $e^+e^- \rightarrow \gamma\gamma$  process is lost. This should be estimated by a detector Monte Carlo, and one can also veto events where a single photon is detected whose recoil photon falls into the insensitive region of the detector with the  $e^+e^- \rightarrow \gamma\gamma$  hypothesis [52]. With the expected energy resolution, Belle-II should, in principle, be able to separate individual peaks of a resonance structure from each other as well as from the background. They could also find the dark photon resonance itself and possibly multiple peaks in lepton pairs, leading to a striking signal, as shown in Figs. 5 and 6.

If a signal is seen, a dedicated low-energy high-intensity  $e^+e^-$  collider would enable more detailed spectroscopy to be performed. From the lattice perspective, it may become possible in the future to compute the vector-vector correlation function for time-like momenta by analytic continuation from the precise lattice calculations [64], in which case one could extract many characteristics of the strongly coupled hidden sector, such as the gauge group, number of colors, number of flavors and the quark masses, by fitting the spectrum.

### 7.3 LHC prospects

Most of the current constraints on the dark photon parameter space from the LHC are expected to improve in the near future, via the search for the Drell-Yan process, search for monojet events, and more detailed studies in  $h \rightarrow 4\ell$  at the LHC Run-2 and beyond.

Projections for the LHC and at a 100 TeV  $pp$  collider are shown in Fig. 7. The 100 TeV Drell-Yan and EWPO LHC14 projections are taken from [30].

Other interesting LHC phenomenology is expected. At the LHC, in the entire allowed parameter space, the dark photon can be produced on-shell and often with a significant amount of  $p_T$ . If the dark photon mass is significantly above the confinement scale, then dark photon decays produce dark quarks. The dark quarks will shower and fragment dominantly into dark mesons. If  $m_\rho/m_\pi \sim 1$ , there is no kinematical suppression for fragmenting into the dark  $\rho$  mesons, and dark  $\rho$  mesons may become the dominant final states due their three spin states. If  $m_\rho < 2m_\pi$ , then for the  $SO(N_c)$  and  $Sp(N_c)$  cases, the singlet dark  $\rho$ -meson can decay into lepton pairs, while the adjoint ones into  $\pi + V^{(*)} \rightarrow \pi + \ell^+ \ell^-$ . The remaining ( $\square\square$  and  $\square$ ) dark  $\rho$ -mesons are stable. Thus, the signals in this case will primarily be a mix of missing energy and  $\rho$ -decays via  $V^{(*)}$  into narrowly collimated small invariant mass lepton pairs, known as ‘lepton jets’ [71–74]. The invariant mass of the lepton pairs could be used to reconstruct the SIMP resonances. The primary challenge in detecting events where  $pp \rightarrow V + X$  is that the lepton jets will be relatively soft because the number of dark hadrons that are typically produced in the fragmentation is large. Depending on the size of the kinetic mixing,  $V$  may decay promptly, or displaced from the collision point. This is an explicit realization of ‘hidden valley’ [49] phenomenology. If the hidden QCD scale (and therefore the dark matter mass) is heavier, emerging jets [75] are possible as well.

Additional states like the  $L = 1$ ,  $S = 1$  dark- $a_1$  state, or even possibly the  $L = 1$ ,  $S = 0$  states, can be produced in the fragmentation of the dark quarks. These decays will produce  $V^{(*)}$  signals. We cannot predict from first principles the rate with which that these high resonances appear in the dark quark jets, though they are likely to be produced at lower rates since they will be heavier than the  $\pi$  and  $\rho$  states. However, if they are produced at significant numbers, then the production of rich lepton-jet events is possible.

## 7.4 ILC prospects

We study the ILC sensitivity using two processes. The first is the search for monophoton events, similar to the search discussed earlier at LEP II with DELPHI data. The second is precision measurements of  $e^+e^- \rightarrow f\bar{f}$  processes, which are usually cast in the language of contact (effective four-fermion) operators.

The monophoton search is limited by the background. We rely on the simulated background levels in Ref. [67], assuming  $\sqrt{s} = 500$  GeV, with integrated luminosity of  $500 \text{ fb}^{-1}$  and beam polarizations  $P_{e^-} = 0.8$ ,  $P_{e^+} = -0.6$ . (This polarization suppresses the background as the electron is nearly right-handed and does not couple to the  $W$  boson.) We use the narrow width approximation, so that the dark photon production appears in one bin of the recoil photon energy, and allow for  $2\sigma$  fluctuations relative to the simulated background. The sensitivity goes far beyond the current limits, as is shown in the green curve in Fig. 7. Unfortunately the simulation in Ref. [67] does not extend to photon energies above  $E_\gamma > 220$  GeV, so the sensitivity for  $m_V < 173$  GeV cannot be evaluated. We hope that studies will be conducted for smaller dark photon masses as well.

Contact operator analyses are the standard analyses in  $e^+e^- \rightarrow f\bar{f}$  processes via the interference with the Standard Model amplitudes, and are expressed by the energy scale of the dimension-six four-fermion operators. Compared to the LEP II limits of approximately 20 TeV [31], the ILC is expected to probe the energy scale up to nearly 200 TeV [68], based on Ref. [69]. Since the center-of-mass energy is fixed at the ILC, unlike the LHC case, it is straightforward to re-interpret the ILC sensitivity to the case where the dark photon mass is lower than the center-of-momentum energy. The resulting projected reach is shown in the violet curve in Fig. 7.

In addition, the ILC will tighten EWPO constraints with its GigaZ option [29, 30], as shown in the solid brown curve in Fig. 7.

### 7.5 Beam dump and fixed target experiments

If the dark photon decays visibly, it can be searched for in beam dump experiments. When a proton beam hits a thick target, many hadrons are produced, some of which may decay into the dark photon. The dark photon may additionally be produced via bremsstrahlung off a proton. After all the hadrons and muons are stopped and/or swept away, only very weakly interacting particles will propagate into the detector. If the dark photon decays into  $\ell^+\ell^-$  or hadrons, it can be detected in this fashion.

Since the dark photon branching fraction into the visible particles is extremely small if it can decay to the strong sector, the sensitivity of beam dump experiments to the dark photon is limited to  $m_V < 2m_\pi$ . However, such a search would be very sensitive to small values of  $\epsilon_\gamma$ . In the orange solid curve of Fig. 7, we show the projected reach of the SHiP (Search for Hidden Particles) facility, proposed for the CERN SPS [66], into the dark photon parameter space.

Finally, the projected reach for the proposed electron fixed-target experiment APEX [70] is shown in the solid dark purple curve of Fig. 7.

## 8 Discussion

In this paper, we discussed the phenomenology of strongly interacting massive particle (SIMP) dark matter. We use a dark photon that is kinetically mixed with  $U(1)_Y$  to maintain kinetic equilibrium with the Standard Model sector during freeze-out. We studied a host of experimental constraints on the dark photon and its coupling to the dark matter. A wide parameter range is allowed by all constraints, summarized in Fig. 4. Stringent limits on dark matter annihilations at the time of recombination as well as from indirect-detection are evaded due to the  $p$ -wave nature of the SIMPs annihilation through the dark photon exchange.

Many current and forthcoming experiments will improve their sensitivity to the remaining viable parameter space, as is summarized in Fig. 7. The novel idea proposed here is to use the dark photon to study the spectroscopy of the vector meson resonances in the SIMP sector. At an  $e^+e^-$  collider, tagging the photon and measuring its energy uniquely determines the recoiling system, even when it is invisible. Alternatively, some resonances may decay back into Standard Model leptons, if their mass is too light to decay into the

SIMP sector. The dark photon peak could be clearly seen in the invariant mass distribution of the mono-photon events, and the resonant structure of the strongly coupled theory could be detected, as is illustrated in Figs. 5 and 6. A single-photon trigger at Belle-II could prove crucial in exploring this dark spectroscopy. The spectroscopy would allow us in principle to determine characteristics of the dark sector such—as the gauge group, the particle content, and the quark masses in the SIMP sector—once a better theoretical handle to compute dynamics of strongly coupled gauge theories is at hand. More generally, such spectroscopy can be used to study any strongly coupled sector that couples to the Standard Model in a conceptually similar way, even if not related to dark matter. At the LHC, the (on-shell or off-shell) dark photon can produce dark quarks which fragment into some of the low-lying resonances which decay into Standard Model leptons or jets, leading to lepton-jet and/or emerging-jet signatures. The direct detection of SIMP dark matter appears possible using semi-conductor detectors with electron recoil signals in parts of the viable parameter space.

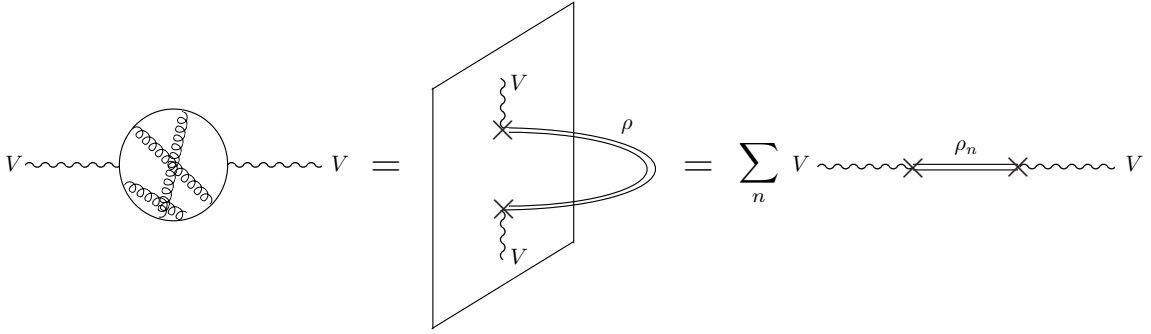
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## A Hadronic widths from holographic QCD

In Section 5.1.2, the main results of this appendix were presented.

We choose to model the resonance structure using the idea of holographic QCD (see Ref. [76] for a review). The idea is as follows. The flavor symmetry in the dark QCD-like sector is promoted to a gauge symmetry in the AdS bulk. A scalar field  $X$  in the bulk, with the same quantum numbers as the quark bilinear condensate, acquires an expectation value and spontaneously breaks the flavor symmetry. The gauge fields have Kaluza–Klein (KK) towers that correspond to the radial excitations of the spin-one bound states. The broken gauge fields are the analog of the  $a_1$  axial-vector mesons of QCD, while the unbroken gauge fields, which also acquire mass due to the boundary conditions [77], are the analog of the  $\rho$



**Figure 8.** Schematic illustration of holographic QCD. The left most diagram represents the vacuum polarization diagram in the strongly coupled QCD-like gauge theory, where the external line is the dark photon. Using holography, it is computed in  $\text{AdS}_5$  where the dark photon lives on the brane while the gauge boson of the flavor symmetry extends into the bulk. Using the KK expansion of the bulk gauge boson, it is replaced by the sum over the KK tower in four dimensions.

vector mesons of QCD. The quantum numbers of the  $\rho$ -mesons can be read off from gauging  $H$ , which matches with the intuition from the quark model with  $L = 0$ ,  $S = 1$  bound states; see Table 2. In the case of  $SU(N_c)$  gauge theories, there is an additional  $U(1)_B$  gauge boson that is the analog of the  $\omega$  vector mesons of QCD. A schematic illustration of the setup is given in Fig. 8. A reader interested directly in the results of the invisible width modeling can skip to the summary given in the final paragraph of this appendix.

Given that we normalize the generators of the flavor symmetry as  $\text{Tr} T^a T^b = 2\delta^{ab}$  [3], as opposed to the more conventional  $\text{Tr} T^a T^b = \frac{1}{2}\delta^{ab}$ , we define the covariant derivative for the bulk gauge field to be

$$D_\mu = \partial_\mu - i\frac{1}{2}g_5 T^a A_\mu \quad (\text{A.1})$$

so that in what follows, the definition of the gauge coupling constant  $g_5$  matches that in the literature. Correspondingly, the Noether currents are

$$J_\mu^a = \frac{1}{2}q^\dagger \gamma_\mu T^a q. \quad (\text{A.2})$$

The spectrum of the unbroken gauge sector is relatively model-independent, fixed only by the boundary conditions of AdS. An early attempt in the literature using the ‘hard wall’ boundary produced a spectrum  $m_n^2 \sim n^2$  [78], which does not match the observed Regge behavior. Instead, we use here the spectrum obtained using the ‘soft wall’ model that reproduces the Regge behavior,  $m_n^2 \propto (n + J)$  [45] (details below). In contrast, the spectrum of the broken gauge sector, namely that of  $a_1$  and  $\pi$  in the case of QCD, is highly model-dependent as it depends on the potential for the field  $X$  in the bulk and its interactions. In the models discussed in this work, the dark photon does not mix into axial vectors, and these are omitted from further discussion in this appendix.

The key point is that in holographic QCD, the sum over the KK towers in the vector-vector two-point function reproduces the parton-level vacuum polarization diagram in the large  $N_c$  limit. Here and below, we only consider the vector currents of the unbroken

symmetries. The definition of the vector-vector correlation function [78] is

$$\int d^4x e^{iq \cdot x} \langle J_\mu^a(x) J_\nu^b(0) \rangle = i \delta^{ab} (q_\mu q_\nu - g_{\mu\nu}) \Pi(q^2). \quad (\text{A.3})$$

Holography states that this is equivalent to the propagation of the gauge boson in the bulk. By writing the gauge boson in terms of its KK states, with the ‘soft-wall’ background in AdS [45], the KK spectrum of the unbroken gauge bosons is simply

$$m_n^2 = 4n, \quad n = 1, 2, \dots \quad (\text{A.4})$$

in units of the AdS curvature. The KK gauge field wave functions are given in terms of the associated Laguerre polynomials,

$$v_{n+1}(z) = z^2 \sqrt{\frac{2n!}{(n+1)!}} L_n^1(z^2). \quad (\text{A.5})$$

Then the vacuum polarization diagram can be obtained by the sum over the KK tower of the gauge bosons,

$$\Pi(q^2) = -\frac{1}{g_5^2} \sum_n \frac{(v_n''(0))^2}{(q^2 - m_n^2 + i\epsilon)m_n^2} = -\frac{1}{g_5^2} \sum_n \frac{2}{q^2 - 4n + i\epsilon}. \quad (\text{A.6})$$

The sum over the tower diverges logarithmically as expected, but is renormalized,

$$\Pi(q^2) - \Pi(0) = -\frac{2}{g_5^2} \sum_n \frac{q^2}{(q^2 - 4n + i\epsilon)4n} = -\frac{2}{g_5^2} \frac{1}{4} H_{-q^2/4}, \quad (\text{A.7})$$

where  $H_r$  is the harmonic number of order  $r$ . Comparing it to Eq. (5.11), we can read off

$$\frac{F_{\rho_n}^2}{m_{\rho_n}^2} = \frac{2}{g_5^2}. \quad (\text{A.8})$$

The overall normalization, determined by  $g_5$ , is still undetermined. Using the asymptotic expansion for large  $-q^2$  in the deep Euclidean region,

$$H_{-q^2/4} \approx -\log \frac{4}{-q^2} + \gamma + O(-q^2)^{-1}, \quad (\text{A.9})$$

the coefficient of the logarithm can be matched to that in the parton-level calculation of the self-energy. The 5D gauge coupling constant in the AdS unit is then determined to be

$$g_5^2 = \frac{12\pi^2}{N_c}. \quad (\text{A.10})$$

The imaginary part of the two-point function gives the decay width of the vector-mesons. With the expressions above, one obtains a series of delta functions with no widths, which would not mimic the realistic situation. In order to properly include widths of the KK gauge bosons, one needs to go beyond the leading approximation in the  $1/N_c$  expansion, which is beyond the current state of holographic QCD. Here, we model the widths by

$U(1)_D \subset H$	embedding	$\rho$ representations
$Sp(N_c)$	$SU(N_f) \times U(1)_D$ $\subset Sp(2N_f)$	$(\square\square, +2) \oplus (\overline{\square\square}, -2) \oplus (\text{adj}, 0) \oplus (1, 0)$
$SO(N_c)$	$SU(N_f/2) \times U(1)_D$ $\subset SO(N_f)$	$(\square, +2) \oplus (\overline{\square}, -2) \oplus (\text{adj}, 0) \oplus (1, 0)$
$SU(N_c)$	$SU(N_1) \times SU(N_2) \times U(1)_D$ $\subset SU(N_f)$	$(\square, \overline{\square}, 2) \oplus (\overline{\square}, \square, -2) \oplus$ $(\text{adj}, 1, 0) \oplus (1, \text{adj}, 0) \oplus (1, 1, 0)$

**Table 2.** Gauging the unbroken flavor symmetry in the AdS bulk,  $H$ , we can identify the spectrum of the  $\rho$ -mesons for the  $Sp(N_c)$ ,  $SO(N_c)$  and  $SU(N_c)$  gauge theories.

the perturbative two-body decays into the lowest KK pion states (the dark matter). The coupling  $g_{\rho\pi\pi}$  is proportional to  $g_5$ , where the proportionality constant is given by the overlap integral of the relevant wave functions in the bulk, which are model-dependent. As a proxy, we simply set  $g_{\rho\pi\pi} = g_5$  which is supported by QCD data ( $g_{\rho\pi\pi} = 5.95$  vs.  $g_5 = (12\pi^2/N_c)^{1/2} = 6.28$  [79]), and obtain

$$\Gamma_n^\rho = \frac{1}{4} D_R \frac{\beta^3}{96\pi} g_5^2 m_n, \quad \beta = \sqrt{1 - \frac{4m_\pi^2}{m_n^2}}. \quad (\text{A.11})$$

Here, the factor  $D_R = \text{Tr}_R (T^a)^2$  (no sum over  $a$ ) is the trace of the generator for the singlet  $\rho$  vector on the representation  $R$  of the pions (including their conjugates, see Table 1). For the gauge groups considered here,

$$D_R = \begin{cases} 4(N_f - 1) & Sp(N_c), \\ 2(N_f + 2) & SO(N_c), \\ 4N_f & SU(N_c). \end{cases} \quad (\text{A.12})$$

The width can be included in the sum (Eq. A.7) by modifying the external momentum

$$q^2 \rightarrow q_{\text{eff}}^2 = q^2 \left( 1 + i\Gamma_n^\rho/m_n(m_n^2 = q^2) \right), \quad (\text{A.13})$$

for  $q^2 > 4m_\pi^2$ . Then the vacuum polarization function is

$$\Pi(q^2) = -\frac{1}{g_5^2} \sum_n \frac{2}{q_{\text{eff}}^2 - 4n + i\epsilon}. \quad (\text{A.14})$$

The inclusion of the width in this manner causes the asymptotic behavior of the harmonic number to change its normalization at  $O(g_5^2)$ , which can be fixed by the rescaling

$$\Pi(q^2) - \Pi(0) = -\frac{N_c}{24\pi^2} \left( 1 - \frac{1}{\pi} \tan^{-1} \frac{1}{4} D_R \frac{g_5^2}{96\pi} \right)^{-1} H_{-q_{\text{eff}}^2/4}. \quad (\text{A.15})$$

Finally, to obtain the decay width for the dark photon  $\Gamma_V$ , the vacuum polarization needs to be normalized for the dark-photon current. The dark photon couples to the

unnormalized generators

$$Q = \begin{cases} \text{diag}(\underbrace{+1, \dots, +1}_{N_f}, \underbrace{-1, \dots, -1}_{N_f}) & Sp(N_c) \\ \text{diag}(\underbrace{+1, \dots, +1}_{N_f/2}, \underbrace{-1, \dots, -1}_{N_f/2}) & SO(N_c) \end{cases} \quad (\text{A.16})$$

where  $Q$  is the unnormalized generator of the dark photon. Therefore the width of the dark photon is given by

$$m_V \Gamma_V = a_V e_D^2 m_V^2 \Im \Pi(m_V^2), \quad (\text{A.17})$$

where  $a_V$  is a normalization factor relating the 2-pt correlation function for the vector  $\rho$ -meson and hidden-photon

$$a_V = \begin{cases} N_f & Sp(N_c), \\ N_f/2 & SO(N_c). \end{cases} \quad (\text{A.18})$$

This is the main result of this section, and Eq. (A.17) will be used to model the width of the dark photon into dark-hadrons.

For  $SU(N_c)$ , the  $SU(N_f)_L \times SU(N_f)_R \times U(1)_B$  symmetry is gauged in the AdS bulk. The  $\rho$ -mesons belong to the unbroken  $SU(N_f)_V$  subgroup, while  $\omega$  belongs to the unbroken  $U(1)_B$ . The vector boson masses of  $\rho$  and  $\omega$  are degenerate because they are both fixed only by the boundary conditions—independent of how the symmetry is broken in the bulk by the  $X$  scalar—in good agreement with QCD. The generator in the vector-vector correlation function is

$$Q = \begin{pmatrix} I_{N_1} & 0 \\ 0 & I_{N_2} \end{pmatrix} = \frac{N_1 - N_2}{N_f} I_{N_f} + \frac{2}{N_f} \begin{pmatrix} N_2 \cdot I_{N_1} & 0 \\ 0 & N_1 \cdot I_{N_2} \end{pmatrix}, \quad (\text{A.19})$$

separated into the  $U(1)_B$  generator and a traceless  $U(1) \subset SU(N_f)$  generator. Therefore the  $\omega$  contribution to the vector-vector two-point function comes with the weight  $(N_1 - N_2)^2/N_f$ , while the  $\rho$  contribution has the weight  $4N_1 N_2/N_f$ . However, pions do not carry baryon number, and hence  $\omega$  does not couple to pions at the ‘tree-level.’ At the quantum level, pions do carry baryon number due to the winding number of its soliton (Skyrmion) solution [6]. Therefore,  $\omega$  decays into  $3\pi$  via the Wess–Zumino–Witten terms, due to interactions of the form

$$\mathcal{L}_{\text{eff}} \sim \frac{g_5^2 N_c}{f_\pi} \epsilon^{\kappa\lambda\mu\nu} \omega_{\kappa\lambda} \rho_{\mu\nu}^a \pi^a + \frac{g_5 N_c}{f_\pi^3} \epsilon^{\kappa\lambda\mu\nu} \omega_{\kappa} \partial_\lambda \pi^a \partial_\mu \pi^b \partial_\nu \pi^c f^{abc}. \quad (\text{A.20})$$

We obtain

$$\begin{aligned} \Pi(q^2) - \Pi(0) = & -\frac{N_c N_f}{12\pi^2} \left[ \frac{4N_1 N_2}{N_f} \left( 1 - \frac{1}{\pi} \tan^{-1} \frac{\Gamma_{\rho_\infty}}{m_{\rho_\infty}} \right)^{-1} H_{-q_{\text{eff},\rho}^2/4} \right. \\ & \left. + \frac{(N_1 - N_2)^2}{N_f} \left( 1 - \frac{1}{\pi} \tan^{-1} \frac{\Gamma_{\omega_\infty}}{m_{\omega_\infty}} \right)^{-1} H_{-q_{\text{eff},\omega}^2/4} \right]. \end{aligned} \quad (\text{A.21})$$



Here,

$$q_{\text{eff},\rho}^2 = q^2 \left( 1 + i \frac{\Gamma_\rho}{m_\rho} (m_\rho^2 = q^2) \right), \quad q_{\text{eff},\omega}^2 = q^2 \left( 1 + i \frac{\Gamma_\omega}{m_\omega} (m_\omega^2 = q^2) \right). \quad (\text{A.22})$$

In summary, the main results of the holographic QCD discussion are as follows: The invisible width of the dark photon into dark-hadrons of the strongly coupled SIMP sector can be modeled in  $Sp(N_c)$  and  $SO(N_c)$  theories using Eq. (A.17), along with Eqs. (A.18), (A.15) and (A.13) and in  $SU(N_c)$  theories using Eqs. (A.17), (A.21) and (A.22) and  $a_V = 1$ .

It would be interesting to study how the phenomenology varies for different gauge groups,  $N_c$ ,  $N_f$ , and quark masses in a comprehensive way. We leave this exploration to future studies.

## References

- [1] Y. Hochberg, E. Kuflik, T. Volansky, and J. G. Wacker, *Mechanism for Thermal Relic Dark Matter of Strongly Interacting Massive Particles*, *Phys. Rev. Lett.* **113** (2014) 171301, [[arXiv:1402.5143](#)].
- [2] E. D. Carlson, M. E. Machacek, and L. J. Hall, *Self-interacting dark matter*, *Astrophys. J.* **398** (1992) 43–52.
- [3] Y. Hochberg, E. Kuflik, H. Murayama, T. Volansky, and J. G. Wacker, *Model for Thermal Relic Dark Matter of Strongly Interacting Massive Particles*, *Phys. Rev. Lett.* **115** (2015), no. 2 021301, [[arXiv:1411.3727](#)].
- [4] J. Wess and B. Zumino, *Consequences of Anomalous Ward Identities*, *Phys. Lett.* **B37** (1971) 95.
- [5] E. Witten, *Global Aspects of Current Algebra*, *Nucl. Phys.* **B223** (1983) 422–432.
- [6] E. Witten, *Current Algebra, Baryons, and Quark Confinement*, *Nucl. Phys.* **B223** (1983) 433–444.
- [7] S.-M. Choi and H. M. Lee, *SIMP Dark Matter with Gauged  $Z_3$  Symmetry*, *JHEP* **09** (2015) 063, [[arXiv:1505.00960](#)].
- [8] H. M. Lee and M.-S. Seo, *Communication with SIMP dark mesons via  $Z$ -portal*, *Phys. Lett.* **B748** (2015) 316–322, [[arXiv:1504.00745](#)].
- [9] N. Bernal, C. García-Cely, and R. Rosenfeld, *WIMP and SIMP Dark Matter from the Spontaneous Breaking of a Global Group*, *JCAP* **1504** (2015), no. 04 012, [[arXiv:1501.01973](#)].
- [10] P. Schwaller, *Gravitational Waves from a Dark Phase Transition*, *Phys. Rev. Lett.* **115** (2015), no. 18 181101, [[arXiv:1504.07263](#)].
- [11] N. Bernal and X. Chu,  *$Z_2$  SIMP Dark Matter*, [[arXiv:1510.08527](#)].
- [12] N. Bernal, X. Chu, C. García-Cely, T. Hambye, and B. Zaldivar, *Production Regimes for Self-Interacting Dark Matter*, [[arXiv:1510.08063](#)].
- [13] M. Hansen, K. Langble, and F. Sannino, *SIMP model at NNLO in chiral perturbation theory*, *Phys. Rev.* **D92** (2015), no. 7 075036, [[arXiv:1507.01590](#)].

- [14] D. Clowe, A. Gonzalez, and M. Markevitch, *Weak Lensing Mass Reconstruction of the Interacting Cluster 1E0657-558: Direct Evidence for the Existence of Dark Matter*, *Astrophys. J.* **604** (2004) 596–603, [[astro-ph/0312273](#)].
- [15] M. Markevitch, A. H. Gonzalez, D. Clowe, A. Vikhlinin, L. David, W. Forman, C. Jones, S. Murray, and W. Tucker, *Direct Constraints on the Dark Matter Self-Interaction Cross-Section from the Merging Galaxy Cluster 1E0657-56*, *Astrophys. J.* **606** (2004) 819–824, [[astro-ph/0309303](#)].
- [16] S. W. Randall, M. Markevitch, D. Clowe, A. H. Gonzalez, and M. Bradac, *Constraints on the Self-Interaction Cross-Section of Dark Matter from Numerical Simulations of the Merging Galaxy Cluster 1E 0657-56*, *Astrophys. J.* **679** (2008) 1173–1180, [[arXiv:0704.0261](#)].
- [17] M. Rocha, A. H. G. Peter, J. S. Bullock, M. Kaplinghat, S. Garrison-Kimmel, J. Onorbe, and L. A. Moustakas, *Cosmological Simulations with Self-Interacting Dark Matter I: Constant Density Cores and Substructure*, *Mon. Not. Roy. Astron. Soc.* **430** (2013) 81–104, [[arXiv:1208.3025](#)].
- [18] A. H. G. Peter, M. Rocha, J. S. Bullock, and M. Kaplinghat, *Cosmological Simulations with Self-Interacting Dark Matter II: Halo Shapes Vs. Observations*, *Mon. Not. Roy. Astron. Soc.* **430** (2013) 105, [[arXiv:1208.3026](#)].
- [19] D. N. Spergel and P. J. Steinhardt, *Observational Evidence for Selfinteracting Cold Dark Matter*, *Phys. Rev. Lett.* **84** (2000) 3760–3763, [[astro-ph/9909386](#)].
- [20] W. J. G. de Blok, *The Core-Cusp Problem*, *Adv. Astron.* **2010** (2010) 789293, [[arXiv:0910.3538](#)].
- [21] M. Boylan-Kolchin, J. S. Bullock, and M. Kaplinghat, *Too Big to Fail? the Puzzling Darkness of Massive Milky Way Subhaloes*, *Mon. Not. Roy. Astron. Soc.* **415** (2011) L40, [[arXiv:1103.0007](#)].
- [22] M. Kaplinghat, S. Tulin, and H.-B. Yu, *Dark Matter Halos as Particle Colliders: a Unified Solution to Small-Scale Structure Puzzles from Dwarfs to Clusters*, [[arXiv:1508.03339](#)].
- [23] J. Zavala, M. Vogelsberger, and M. G. Walker, *Constraining Self-Interacting Dark Matter with the Milky Way’s Dwarf Spheroidals*, *Monthly Notices of the Royal Astronomical Society: Letters* **431** (2013) L20–L24, [[arXiv:1211.6426](#)].
- [24] M. Vogelsberger, J. Zavala, and A. Loeb, *Subhaloes in Self-Interacting Galactic Dark Matter Haloes*, *Mon. Not. Roy. Astron. Soc.* **423** (2012) 3740, [[arXiv:1201.5892](#)].
- [25] R. Massey et al., *The behaviour of dark matter associated with four bright cluster galaxies in the 10 kpc core of Abell 3827*, *Mon. Not. Roy. Astron. Soc.* **449** (2015), no. 4 3393–3406, [[arXiv:1504.03388](#)].
- [26] F. Kahlhoefer, K. Schmidt-Hoberg, J. Kummer, and S. Sarkar, *On the Interpretation of Dark Matter Self-Interactions in Abell 3827*, *Mon. Not. Roy. Astron. Soc.* **452** (2015), no. 1 L54–L58, [[arXiv:1504.06576](#)].
- [27] A. Hook, E. Izaguirre, and J. G. Wacker, *Model Independent Bounds on Kinetic Mixing*, *Adv. High Energy Phys.* **2011** (2011) 859762, [[arXiv:1006.0973](#)].
- [28] E. Izaguirre, G. Krnjaic, P. Schuster, and N. Toro, *New Electron Beam-Dump Experiments to Search for MeV to Few-GeV Dark Matter*, *Phys. Rev.* **D88** (2013) 114015, [[arXiv:1307.6554](#)].

- [29] R. Essig, J. Mardon, M. Papucci, T. Volansky, and Y.-M. Zhong, *Constraining Light Dark Matter with Low-Energy  $e^+e^-$  Colliders*, *JHEP* **11** (2013) 167, [[arXiv:1309.5084](#)].
- [30] D. Curtin, R. Essig, S. Gori, and J. Shelton, *Illuminating Dark Photons with High-Energy Colliders*, *JHEP* **02** (2015) 157, [[arXiv:1412.0018](#)].
- [31] **DELPHI, OPAL, LEP Electroweak, ALEPH, L3** Collaboration, S. Schael et al., *Electroweak Measurements in Electron-Positron Collisions at W-Boson-Pair Energies at LEP*, *Phys. Rept.* **532** (2013) 119–244, [[arXiv:1302.3415](#)].
- [32] H. Baer, T. Barklow, K. Fujii, Y. Gao, A. Hoang, S. Kanemura, J. List, H. E. Logan, A. Nomerotski, M. Perelstein, et al., *The International Linear Collider Technical Design Report - Volume 2: Physics*, [arXiv:1306.6352](#).
- [33] **BaBar** Collaboration, J. P. Lees et al., *Search for a Dark Photon in  $e^+e^-$  Collisions at BaBar*, *Phys. Rev. Lett.* **113** (2014), no. 20 201801, [[arXiv:1406.2980](#)].
- [34] **CMS** Collaboration, *Properties of the Higgs-like boson in the decay  $H$  to  $ZZ$  to  $4l$  in  $pp$  collisions at  $\sqrt{s} = 7$  and  $8$  TeV*, *CMS-PAS-HIG-13-002*, .
- [35] **CMS** Collaboration, S. Chatrchyan et al., *Measurement of the differential and double-differential Drell-Yan cross sections in proton-proton collisions at  $\sqrt{s} = 7$  TeV*, *JHEP* **12** (2013) 030, [[arXiv:1310.7291](#)].
- [36] J. M. Cline, G. Dupuis, Z. Liu, and W. Xue, *The Windows for Kinetically Mixed  $Z'$ -Mediated Dark Matter and the Galactic Center Gamma Ray Excess*, *JHEP* **08** (2014) 131, [[arXiv:1405.7691](#)].
- [37] **ATLAS** Collaboration, *Search for high-mass dilepton resonances in  $20\text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 8$  TeV with the ATLAS experiment*, *ATLAS-CONF-2013-017*, .
- [38] **BaBar** Collaboration, B. Aubert et al., *Search for Invisible Decays of a Light Scalar in Radiative Transitions  $\Upsilon(3S) \rightarrow \gamma A^0$* , in *Proceedings, 34th International Conference on High Energy Physics (ICHEP 2008)*, 2008. [arXiv:0808.0017](#).
- [39] I. M. Shoemaker and L. Vecchi, *Unitarity and Monojet Bounds on Models for DAMA, CoGeNT, and CRESST-II*, *Phys. Rev.* **D86** (2012) 015023, [[arXiv:1112.5457](#)].
- [40] **DELPHI** Collaboration, J. Abdallah et al., *Search for one large extra dimension with the DELPHI detector at LEP*, *Eur. Phys. J.* **C60** (2009) 17–23, [[arXiv:0901.4486](#)].
- [41] S. L. Adler and W. A. Bardeen, *Absence of Higher Order Corrections in the Anomalous Axial Vector Divergence Equation*, *Phys. Rev.* **182** (1969) 1517–1536.
- [42] E. Kuflik, M. Perelstein, N. R.-L. Lorier, and Y.-D. Tsai, *Elastically Decoupling Dark Matter*, [arXiv:1512.04545](#).
- [43] M. S. Madhavacheril, N. Sehgal, and T. R. Slatyer, *Current Dark Matter Annihilation Constraints from CMB and Low-Redshift Data*, *Phys. Rev.* **D89** (2014) 103508, [[arXiv:1310.3815](#)].
- [44] R. Essig, E. Kuflik, S. D. McDermott, T. Volansky, and K. M. Zurek, *Constraining Light Dark Matter with Diffuse X-Ray and Gamma-Ray Observations*, *JHEP* **11** (2013) 193, [[arXiv:1309.4091](#)].
- [45] A. Karch, E. Katz, D. T. Son, and M. A. Stephanov, *Linear Confinement and  $AdS/QCD$* , *Phys. Rev.* **D74** (2006) 015005, [[hep-ph/0602229](#)].
- [46] G. 't Hooft, *A Planar Diagram Theory for Strong Interactions*, *Nucl. Phys.* **B72** (1974) 461.

- [47] B. Batell, M. Pospelov, and A. Ritz, *Probing a Secluded  $U(1)$  at B-Factories*, *Phys. Rev. D* **79** (2009) 115008, [[arXiv:0903.0363](#)].
- [48] **Particle Data Group** Collaboration, K. A. Olive et al., *Review of Particle Physics*, *Chin. Phys. C* **38** (2014) 090001.
- [49] M. J. Strassler and K. M. Zurek, *Echoes of a Hidden Valley at Hadron Colliders*, *Phys. Lett. B* **651** (2007) 374–379, [[hep-ph/0604261](#)].
- [50] J. Kang and M. A. Luty, *Macroscopic Strings and ‘Quirks’ at Colliders*, *JHEP* **11** (2009) 065, [[arXiv:0805.4642](#)].
- [51] F. A. Berends, G. J. H. Burgers, C. Mana, M. Martínez, and W. L. van Neerven, *Radiative Corrections to the Process  $e^+e^-$  to Neutrino Anti-Neutrino  $\gamma$* , *Nucl. Phys. B* **301** (1988) 583.
- [52] C. Hearty, T. Higuchi, Y. Iwasaki, T. Iwashita, C. Li, and K. Miyabayashi. private communications.
- [53] **SuperCDMS** Collaboration, R. Agnese et al., *Search for Low-Mass Weakly Interacting Massive Particles Using Voltage-Assisted Calorimetric Ionization Detection in the SuperCDMS Experiment*, *Phys. Rev. Lett.* **112** (2014), no. 4 041302, [[arXiv:1309.3259](#)].
- [54] **SuperCDMS** Collaboration, R. Agnese et al., *Search for Low-Mass Weakly Interacting Massive Particles with SuperCDMS*, *Phys. Rev. Lett.* **112** (2014), no. 24 241302, [[arXiv:1402.7137](#)].
- [55] **SuperCDMS** Collaboration, R. Agnese et al., *New Results from the Search for Low-Mass Weakly Interacting Massive Particles with the CDMS Low Ionization Threshold Experiment*, *Phys. Rev. Lett.* **116** (2016), no. 7 071301, [[arXiv:1509.02448](#)].
- [56] R. Essig, J. Mardon, and T. Volansky, *Direct Detection of Sub-GeV Dark Matter*, *Phys. Rev. D* **85** (2012) 076007, [[arXiv:1108.5383](#)].
- [57] P. W. Graham, D. E. Kaplan, S. Rajendran, and M. T. Walters, *Semiconductor Probes of Light Dark Matter*, *Phys. Dark Univ.* **1** (2012) 32–49, [[arXiv:1203.2531](#)].
- [58] Y. Hochberg, Y. Zhao, and K. M. Zurek, *Superconducting Detectors for Super Light Dark Matter*, [[arXiv:1504.07237](#)].
- [59] Y. Hochberg, M. Pyle, Y. Zhao, and K. M. Zurek, *Detecting Superlight Dark Matter with Fermi-Degenerate Materials*, [[arXiv:1512.04533](#)].
- [60] R. Essig, A. Manalaysay, J. Mardon, P. Sorensen, and T. Volansky, *First Direct Detection Limits on Sub-GeV Dark Matter from XENON10*, *Phys. Rev. Lett.* **109** (2012) 021301, [[arXiv:1206.2644](#)].
- [61] **XENON10** Collaboration, J. Angle et al., *A search for light dark matter in XENON10 data*, *Phys. Rev. Lett.* **107** (2011) 051301, [[arXiv:1104.3088](#)]. [Erratum: *Phys. Rev. Lett.* **110**, 249901(2013)].
- [62] R. Essig, M. Fernandez-Serra, J. Mardon, A. Soto, T. Volansky, and T.-T. Yu, *Direct Detection of Sub-GeV Dark Matter with Semiconductor Targets*, [[arXiv:1509.01598](#)].
- [63] **Belle, BaBar** Collaboration, A. J. Bevan et al., *The Physics of the B Factories*, *Eur. Phys. J. C* **74** (2014) 3026, [[arXiv:1406.6311](#)].
- [64] Z. Fodor. private communications.

- [65] A. Soffer, *Constraints on dark forces from the B factories and low-energy experiments*, in *Interplay between Particle and Astroparticle physics London, United Kingdom, August 18-22, 2014*, 2014. [arXiv:1409.5263](#).
- [66] S. Alekhin et al., *A facility to Search for Hidden Particles at the CERN SPS: the SHiP physics case*, [arXiv:1504.04855](#).
- [67] A. Chaus, J. List, and M. Titov, *Model-independent WIMP searches at ILC with single photon*, . a talk presented at International Workshop on Future Linear Colliders, Belgrade, 06–10 October 2014, <http://agenda.linearcollider.org/event/6389/session/1/contribution/243/material/slides/1.pdf>.
- [68] **ILC** Collaboration, G. Aarons et al., *International Linear Collider Reference Design Report Volume 2: Physics at the ILC*, [arXiv:0709.1893](#).
- [69] S. Riemann, *Fermion pair production at a linear collider – a sensitive tool for new physics searches*, . [http://tesla.desy.de/new\\_pages/TDR\\_CD/PartIII/references/LC-TH-2001-007.pdf](http://tesla.desy.de/new_pages/TDR_CD/PartIII/references/LC-TH-2001-007.pdf).
- [70] R. Essig, P. Schuster, N. Toro, and B. Wojtsekhowski, *An Electron Fixed Target Experiment to Search for a New Vector Boson  $A'$  Decaying to  $e^+e^-$* , *JHEP* **02** (2011) 009, [[arXiv:1001.2557](#)].
- [71] N. Arkani-Hamed and N. Weiner, *LHC Signals for a Superunified Theory of Dark Matter*, *JHEP* **12** (2008) 104, [[arXiv:0810.0714](#)].
- [72] M. Baumgart, C. Cheung, J. T. Ruderman, L.-T. Wang, and I. Yavin, *Non-Abelian Dark Sectors and Their Collider Signatures*, *JHEP* **04** (2009) 014, [[arXiv:0901.0283](#)].
- [73] Y. Bai and Z. Han, *Measuring the Dark Force at the LHC*, *Phys. Rev. Lett.* **103** (2009) 051801, [[arXiv:0902.0006](#)].
- [74] C. Cheung, J. T. Ruderman, L.-T. Wang, and I. Yavin, *Lepton Jets in (Supersymmetric) Electroweak Processes*, *JHEP* **04** (2010) 116, [[arXiv:0909.0290](#)].
- [75] P. Schwaller, D. Stolarski, and A. Weiler, *Emerging Jets*, *JHEP* **05** (2015) 059, [[arXiv:1502.05409](#)].
- [76] J. Erlich, *An Introduction to Holographic QCD for Nonspecialists*, *Contemp. Phys.* **56** (2015) 159, [[arXiv:1407.5002](#)].
- [77] C. Csaki, C. Grojean, H. Murayama, L. Pilo, and J. Terning, *Gauge Theories on an Interval: Unitarity without a Higgs*, *Phys. Rev.* **D69** (2004) 055006, [[hep-ph/0305237](#)].
- [78] J. Erlich, E. Katz, D. T. Son, and M. A. Stephanov, *QCD and a Holographic Model of Hadrons*, *Phys. Rev. Lett.* **95** (2005) 261602, [[hep-ph/0501128](#)].
- [79] D. Garcia Gudino and G. Toledo Sanchez, *The  $\omega\rho\pi$  Coupling and the Influence of Heavier Resonances*, *J. Phys. Conf. Ser.* **378** (2012) 012040.